LIGO Signals from the Mirror World





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1. PARITY VIOLATION MIRROR AND **PARTICLES**

of the Talk

Plan 2. GRAVITATIONAL WAVES: THEORY AND DETECTION

> 3. UNEXPECTED LIGO **EVENTS**

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Binary Objects Creation Mechanisms

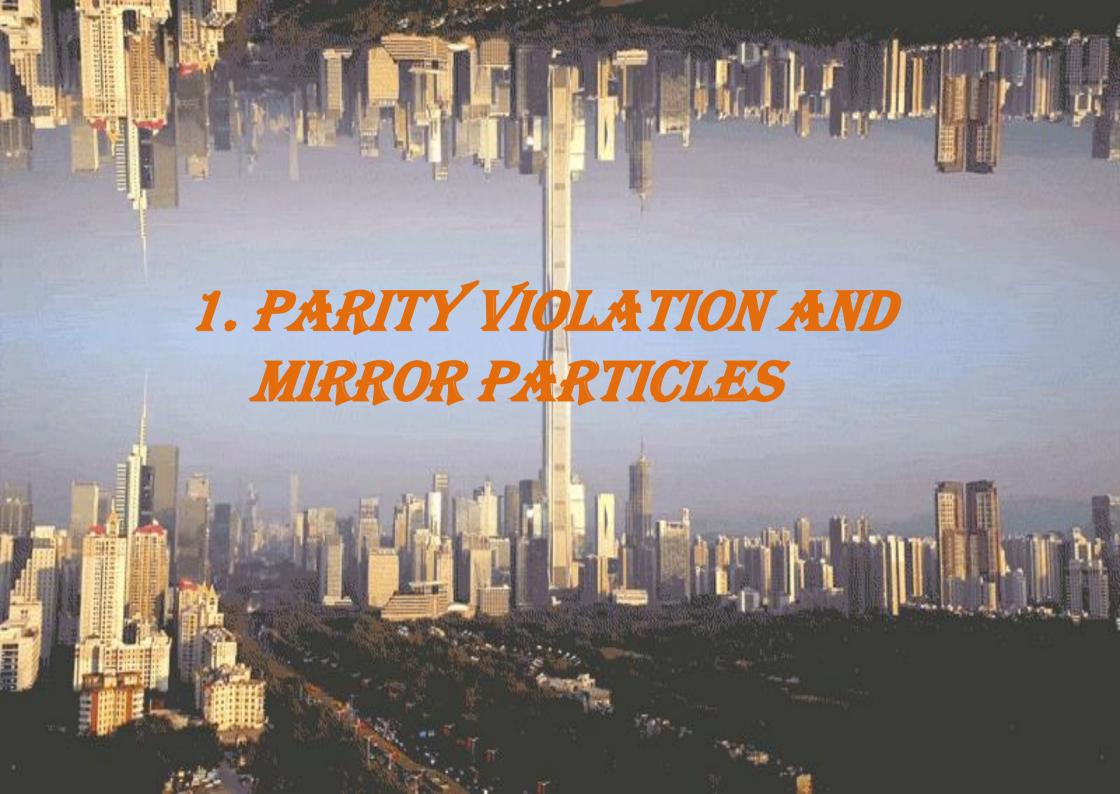
Theoretical BBH Merger Rate

Problems with LIGO Data

In Mirror World

Star Formation Rate

LIGO Signals from Mirror World

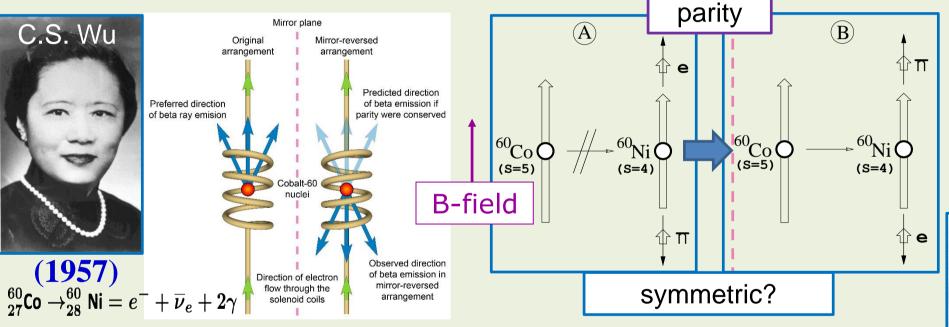


Parity Violation

 τ - θ puzzle: Two different decays were found for charged strange mesons:

$$\theta^+ \rightarrow \pi^+ \pi^0$$
 parity = +1
 $\tau^+ \rightarrow \pi^+ \pi^+ \pi^-$ parity = -1

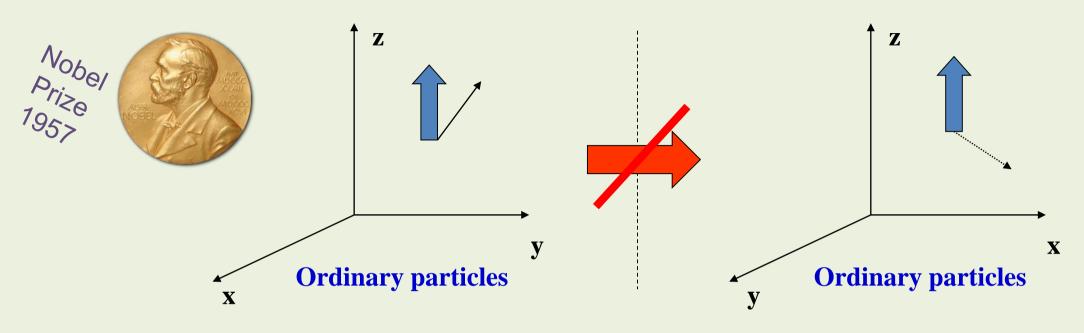
Both reactions represent decay of the same particle, K^+ , with violation of parity.



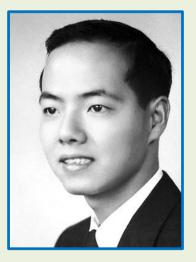
"I cannot believe God is a weak left-hander" Wolfgang Pauli



Mirror Asymmetry



T.D. Lee



C.N. Yang



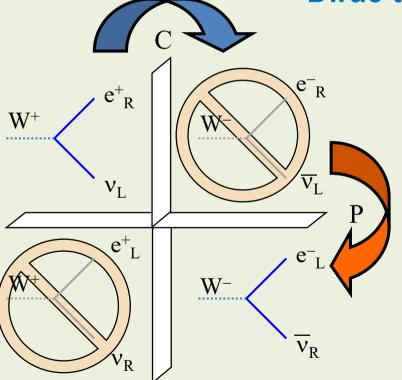
Parity (P) violation means non-equivalence of Left- and Right-handed coordinate systems. To restore the symmetry, reflection should be generalized. Together with P-transformations, particles should be changed by their mirror partners (Lee, Yang - 1956).

CP-violation and Mirror Symmetry



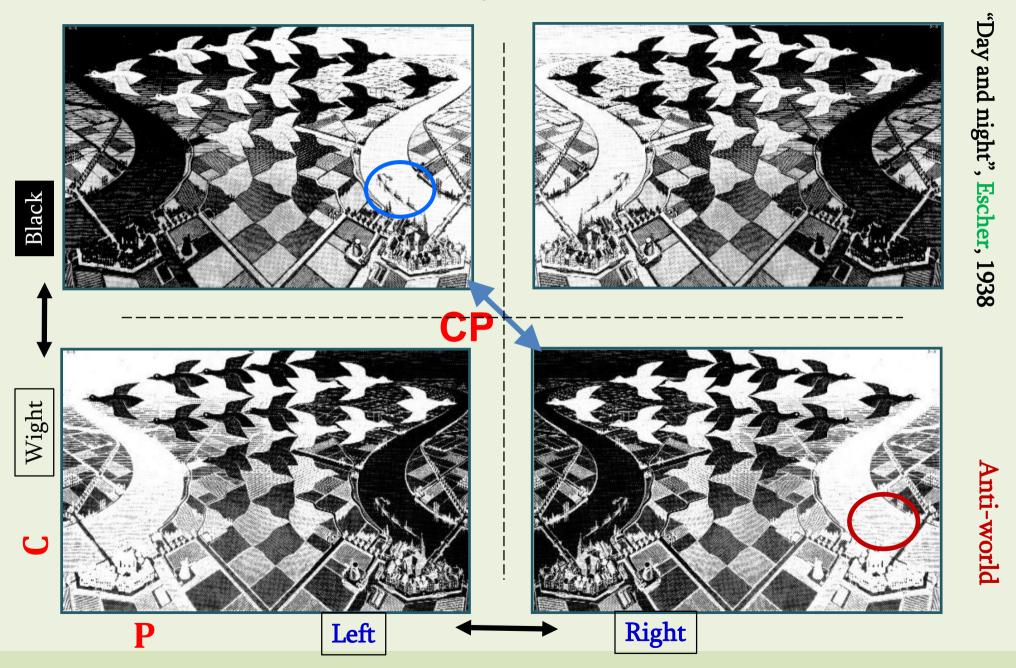
Lee, Landau,... (1957) an economic solution: CP-invariance assumes that antiparticles play the role of mirror partners. Then non-conservation of parity can be introduced without assuming asymmetry of space with respect to inversion.

Dirac theory is symmetric under CP transformations!



Discovery of CP-violation in certain types of weak decay (1964) put again the question of proper choice for the set of mirror partners.

CP-transformations



HIP - 26 Sep 2023



Mirror Partners

Kobzarev, Okun, Pomeranchuk (1966)

$$(\mathbf{0}) \equiv \begin{pmatrix} v \\ e \\ p \\ n \end{pmatrix}_{o}$$

$$(\mathbf{M}) \equiv \begin{pmatrix} v \\ e \\ p \\ n \end{pmatrix}_{M}$$

$$(\mathbf{M}) = \begin{pmatrix} v \\ e \\ p \\ n \end{pmatrix}_{M}$$

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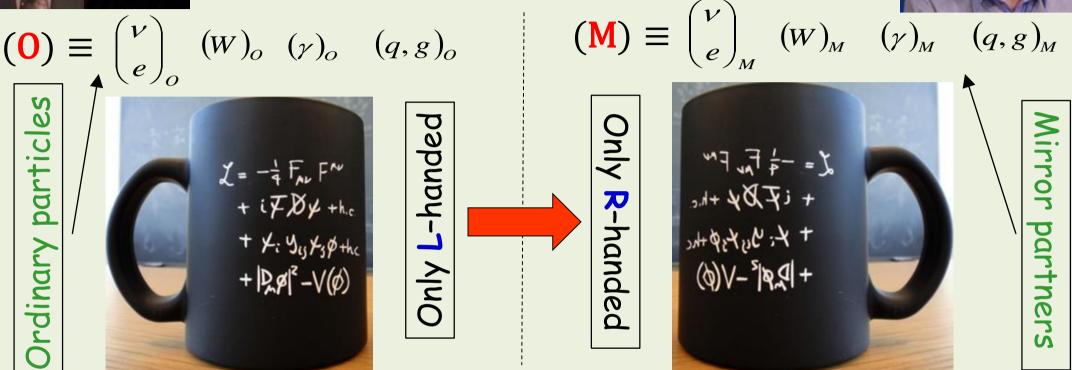
The equivalence between L and R systems is restored if a hidden mirror sector exists in which parity is violated in the opposite way and reflection is accompanied by change of ordinary particles by their mirror partners.

If mirror partners are strictly symmetric to ordinary particles and they can not have usual electromagnetic, weak and strong interactions (doubling of atomic levels, pion states, etc.).



Mirror World

Blinnikov, Khlopov, Berezhiani (1980-90s)



Mirror matter, also called Shadow matter or Alice matter, is a hypothetical counterpart to ordinary matter. One needs to suppose that there is no (or very weak) common interaction between ordinary particles and their mirror partners, except of gravity. Initially it was assumed that all initial conditions, masses and coupling constants of mirror particles were strictly symmetric to ordinary ones.

If our Universe = SM + SM'

Everything can be explained by SM, but there could be more than one SM!

L - R symmetric SM'

L - R symmetric





 $G = G_{SM} \times G'_{SM} = [SU(3)_c \times SU(2)_L \times U(1)_Y] \times [SU(3)_c \times SU(2)_R \times U(1)_Y]$

 $L = L_{SM} + L'_{SM} + L_{int}$ is symmetric under parity exchange: $L_{SM} \leftrightarrow L'_{SM}$



The **Pati-Salam** model (the first **GUT - 1974**) is L - R symmetric. Predicts a high energy **R-**handed weak interactions with heavy **W'** and **Z'** bosons and **R-**neutrinos.

Gauge symmetry group: $(SU(4) \times SU(2)_L \times SU(2)_R)/Z_2$

Motivations

Elegance suggests exact parity conservation → MIRROR MATTER

Historic precedence: Lorentz invariance ↔ antiparticles

Parity symmetry \leftrightarrow mirror particles ?

Mirror matter can explain several physical phenomena:

In astrophysics:

- MACHOs;
- Isolated planets and stars;
- Drag force on Pioneer 10, 11;
- Missing comets;
- Mirror bodies bombardment.

In cosmology:

- ***** Viable DM Candidate;
- * High energy cosmic events.

Higgs – mirror Higgs mixings;

In particle physics:

- > Neutrino sterile neutrino mixings;
- Photon-mirror photon kinetic mixing;
- > Anomalous disappearance of neutrons.

Yin-Yang duality

(in Chinese, Dark-Bright)

Opposite forces are complementary, connected and interdependent



Shadow Particles

Properties:

- > Feebly coupled to SM particles and have the same masses;
- **▶** Interact with each other with shadow (SM') interactions;
- > Stable, singlets under SM;
- > Scarce at the scale of the solar system but could dominate on a larger scale.

Can interact with us via:

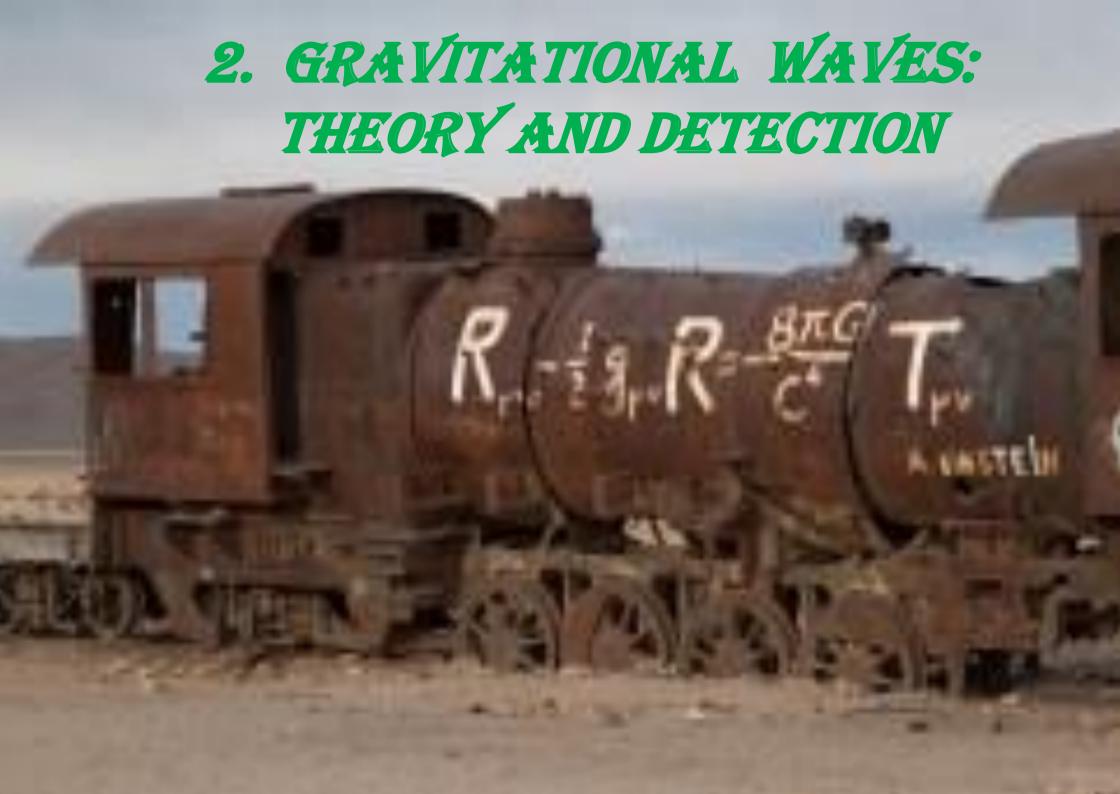
- ☐ Gravity;
- **☐** Mixing with ordinary particles in the neutral sectors.

Problems:

- Symmetric initial conditions \implies equal M and O baryonic densities;
- Mirror particles double species in BBN \implies He abundance > 28%.

Solution:

The temperature of the mirror sector is a few times smaller!



Theory of GWs

Gravitational Wave (**GW**) solution of linearized **Einstein** equations correspond to space-time fluctuations that propagate through the Universe.

The main stages in development of the **GW** theory are:

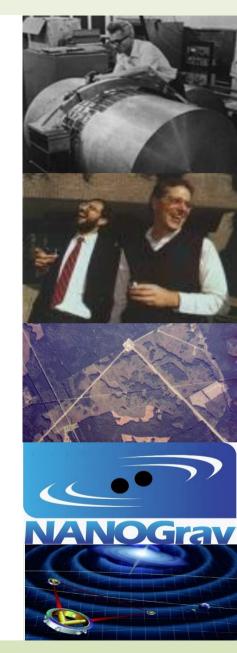
- ☐ In 1916 Albert Einstein claimed that gravitational waves were an inevitable consequence of the linearized General Relativity.
- ☐ In 1936 Einstein, together with Nathan Rosen, derived the opposite result and submitted the paper "Do gravitational waves exist?" to Physical Review.
- On semi-centennial conference in **Bern** (1955) Rosen showed that the energy-momentum pseudo-tensor for **Einstein-Rosen** waves is vanishing.
- On Chapel Hill conference (1956) Richard Feynman proposed a "sticky-bead" experiment to show that GWs carry energy.



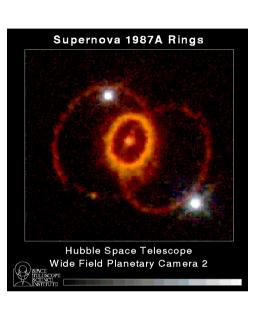
GWs Detection Efforts

The main stages in the **GW** detection are:

- ☐ First attempts in **1960**s by **Weber** using aluminum resonant bar detectors;
- ☐ In 1963 Gertsenshtein and Pustovoit suggested Michelson-Morley type interferometers;
- □ In 1974 Taylor and Hulse detected the binary pulsar PSR B1913+16. Later, some more relativistic binary pulsars were discovered. Orbital period decay in all cases is consistent with the energy loss predicted by General Relativity. Nobel Prize in 1993 for the first observational evidence for GW;
- On 11 February 2016, the LIGO/Virgo collaboration announced the first direct observation of GWs in September 2015. The 2017 Nobel Prize in Physics Barish, Thorne and Weiss;
- ☐ On 28 June 2023, the NANOGrav collaboration announced evidence for a GW stochastic background using observational data from an array of millisecond pulsars.
- ☐ The space missions: **LISA** (**ESA**, to launch in **2037**?) and **DECIGO** (Japan, to launch in **2027**?) were proposed.



What Might Make GWs

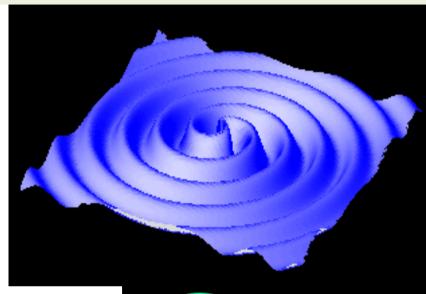


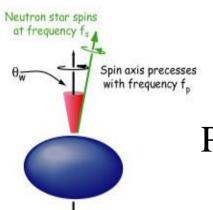
Compact binary inspiral:

"chirps"

Supernovae:

"bursts"



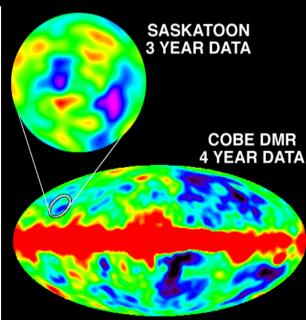


Cosmological Signal:

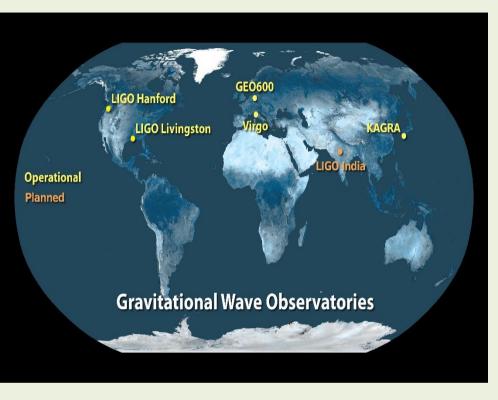
"stochastic background"

Pulsars in our galaxy:

"periodic"



LIGO Observatory



LIGO - gravitational wave detectors in

Hanford and
Livingston,
4 km tunnels,
separation 3,000 km.
VIRGO - in
Cascina, Italy
KAGRA - in
Kamioka, Japan



New Era of Multi-Messenger Astrophysics!

"Observation of GWs from a Binary Black Hole Merger" LIGO/Virgo, Phys. Rev. Lett. 116, 061102 (11 February 2016)

"GW170817: Observation of GWs from a Binary Neutron Star Inspiral" LIGO/Virgo, Phys. Rev. Lett. 119, 161101 (16 October 2017)

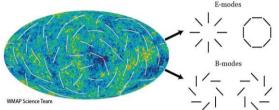




Present Situation

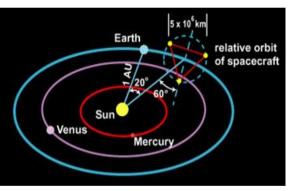
Intensive efforts to find burst, binary coalescence, continuous wave, stochastic

GWs in coincidence with **EM** and neutrino signals;

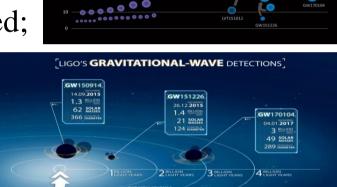


Stochastic background of relict GWs probably was observed in 2023;

The universe has more stellar mass BHs than expected;

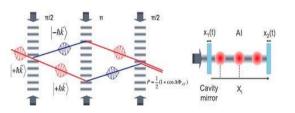


LISA (ESA): three Interferometers 2.5×10⁶ km arm lengths. Planned launch 2037 (earlier launch 2034?);



Black Holes of Known Mass

First **KAGRA** detections with **LIGO/Virgo** were reported on 11 November 2021. The O4 Observing Run was started on 25 May 2023.



3rd generation detectors: **Einstein Telescope** (EU) and **Cosmic Explorer** (US), with atom interferometers in **0.1–10 Hz** (between **LISA** and **LIGO-Virgo-KAGRA**).

Events Detected So Far

After the analysis of first three observing runs

O1, O2, O3a & O3b

there are 90 events with a probability of

 $P_{astro} > 0.5$

being of astrophysical origin.

"GWTC-3: Compact Binary Coalescences Observed by LIGO and Virgo During the Second Part of the Third Observing Run"

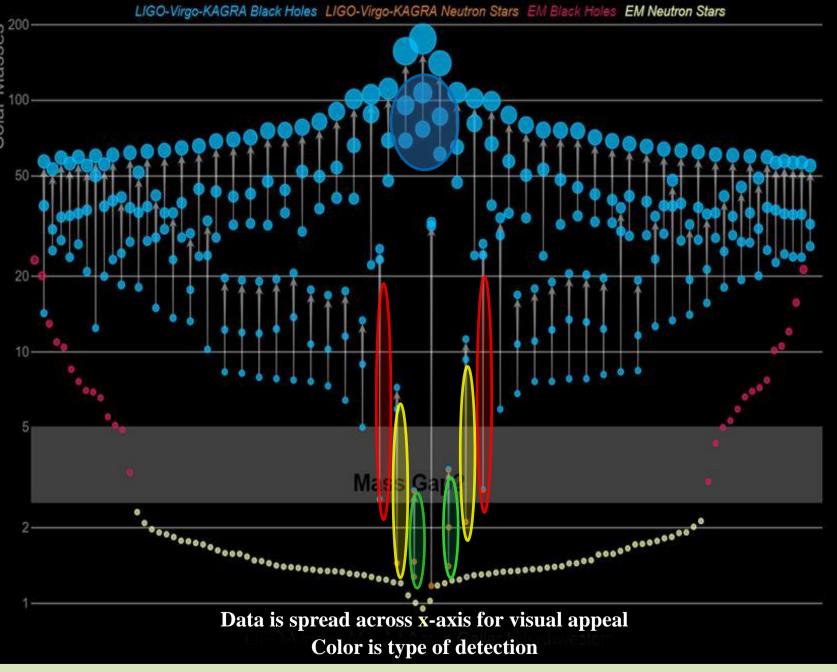
LIGO & VIRGO & KAGRA, arXiv: 2111.03606

Merger objects:	BH-BH	NS-NS	BH-NS	BH-Mass gap
Number of events:	84	2	2	2

Only one NS-NS merger had an accompanying Electromagnetic counterpart!



Masses in the Stellar Graveyard



NS-NS

BH-NS

BH-lower mass gap

Upper mass gap, intermediate mass **BH**

Interactive figure:
https://ligo.northwestern.
edu/media/mass-
plot/index.html

Unexpected Events

➤ GW190521 & GW190426_190642: First ever observation of intermediate mass BHs

$$95M_{\odot} - 69M_{\odot} \rightarrow 156M_{\odot}$$

 $107M_{\odot} - 77M_{\odot} \rightarrow 175M_{\odot}$

Many models of star evolution predict existence of upper mass gap $65M_{\odot} - 135M_{\odot}$ for remnant compact objects.

He core mass	Process	Remnant compact object	
$32-65M_{\odot}$	Pulsational Pair-instability	≤ 65M _O	
$65 - 135 M_{\odot}$	Pair-instability (e+e- pairs)	Explodes – no remnant	
≥ 135M _O	Direct collapse into BH	≥ 135 <i>M</i> _⊙	

Hierarchical Mergers

Merger rate of **GW190521**-like events:

$$f_{exp} = 0.02 - 0.43 \text{ Gpc}^{-3} yr^{-1}$$

LIGO /Virgo, ApJ. Lett. 900 (2020) L13

Theoretical estimate:

$$f = f_{1g} \times f_{triple} \times f_{survival} \times f_{merger}$$

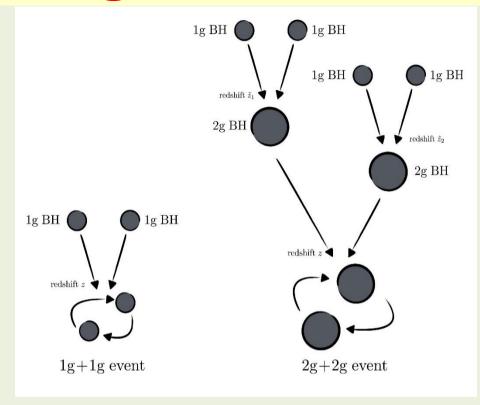
Liu & Lai, MNRAS 502 (2021) 2049

$$f_{1g} \sim (10 - 100) \text{ Gpc}^{-3} yr^{-1}$$

LIGO /Virgo, Phys. Rev. X 9 (2019) 031040

$$f_{triple} \simeq 50 \%$$
 $f_{survival} \simeq 60 \%$

$$f_{theor} = 0.6 - 6 \,\mathrm{Gp}c^{-3}yr^{-1}$$



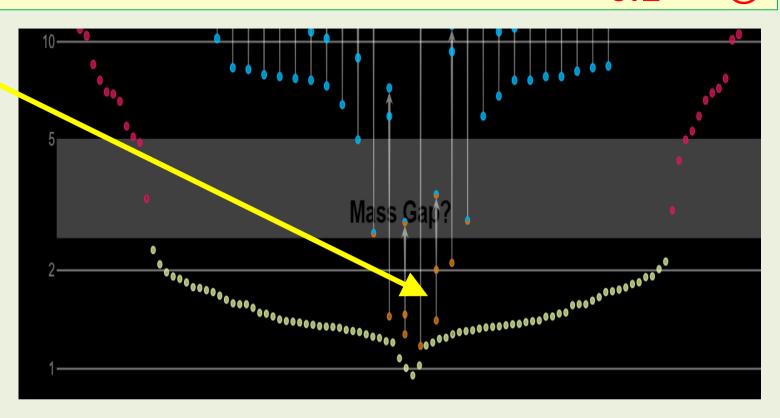
Gerosa. & Berti, Phys. Rev. D 95 (2017)

$$f_{merger} \simeq 20 \%$$

In price of extremal assumptions!

GW190425: Most Massive BNS $3.4^{+0.3}_{-0.1}~M_{\odot}$

X-ray binaries (normal star and a collapsed star) give the lower mass gap $2-5M_{\odot}$ for BHs!

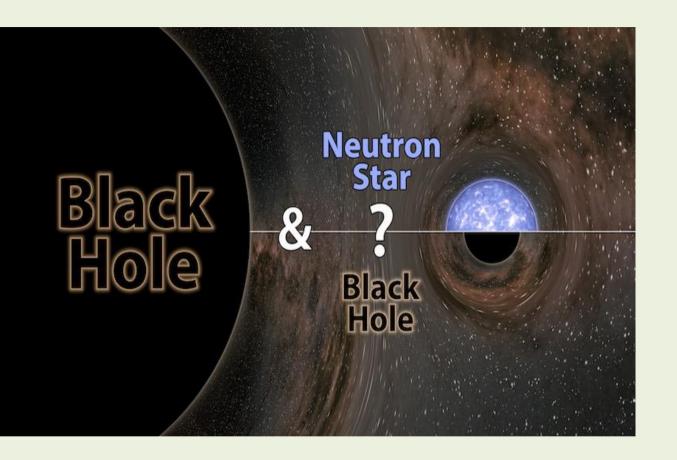


➤ **GW190425-**like systems could be obtained as a result of evolution of ultra-tight binary **He**-star – **NS** systems;

LIGO/Virgo, ApJ Lett. 892 (2020) L3

A phase of mass transfer from a post-helium main-sequence star onto NS is required.

GW190814&GW200210_092254: BH Mass-gap



First components are clearly **BH**s.

Origins of second components with masses $2.59^{+0.08}_{-0.09}M_{\odot}$ &2. $83^{+0.48}_{-0.43}M_{\odot}$ are controversial.

They are heavier than any known pulsars, and lighter than any known BHs so far.



4. WORLD IS RULED BY DARK FORCES

Binary Objects Creation Mechanisms

1. Primordial Black Holes (Sasaki, Suyama, Tanaka & Yokoyama, 2018)
PBH abundance is constrained by microlensing, CMB spectral distortion and wide binaries.

2. Astrophysical binary systems:

- Common envelope evolution; (Giacobbo & Mapelli, 2018)
- ➤ Chemically homogenous evolution; (Mandel & de Mink. 2016)
- > Dynamical processes in dense stellar clusters. (Askar, et al. 2017)

Main formation mechanisms predict low binary merging rates:

$$\mathcal{R}_{theor}^{BBH} \simeq 5 - 10 \text{ Gpc}^{-3}yr^{-1} < \mathcal{R}_{LIGO}^{BBH} = 17 - 45 \text{ Gpc}^{-3}yr^{-1}$$

$$LIGO \& VIRGO \& KAGRA, arXiv: 2111.03634$$

Theoretical BBH Merger Rate

$$\mathcal{R} = \frac{1}{2} \varepsilon P(\tau) N_{BH}$$
 $\varepsilon \simeq 0.01 - 0.001$ - Efficiency coefficient $P(\tau)$ - Delay time distribution (time to merging)

Number of Black Holes: (Elbert, Bullock & Kaplinghat, 2018)

$$N_{BH} = SFR(z) \times \int \varphi(m)N(m) \int f(Z,m) \int \xi(M)dMdZdm$$

 $\varphi(m)$ - Galactic mass function $\xi(M)$ - Initial mass function

$$N(m) = \frac{m}{\int M\xi(M)dM}$$

 $N(m) = \frac{m}{\int M\xi(M)dM}$ - Number of stars in the galaxy of mass m

f(Z, m) - Metallicity distribution function

Star formation rate: (Madau & Dickinson, 2014)

$$SFR(z) = 0.015 \frac{(1+z)^{2.7}}{1 + [(1+z)/2.9]^{5.6}} M_{\odot} Mpc^{-3}yr^{-1}$$

Peaks at: $z \sim 2 \approx t_{lookback} \sim 10.3 \text{ Gyr}$

Problems with LIGO Data

- > Observed Merger Rates are higher than theoretical predictions
- > Only 1 out of 90 events had EM counterpart, while:
 - BNS merger must always be accompanied by Gamma-Ray Bursts;
 - BH-NS mergers in many configurations should emit EM-radiation;
 - If BHs accrete matter they can also emit EM-radiation.
- Mass gap events observed

Suggestion:

GWs detected by LIGO may be emitted by the Mirror World binaries!

In Mirror World

M-world, along with O-world, was created by $Big\ Bang$, but with lower reheating T.

- \triangleright Constrain from Big Bang Nucleosynthesis: $x \equiv T'/_T < 0.64$
- ightharpoonup Certain leptogenesis mechanism gives: $1 \le \frac{n_b}{n_b} \lesssim 10$
- > Mirror World can explain all Dark Matter: $\frac{\Omega'_b}{\Omega_b} \approx 5$
- ➤ Helium abundance is higher: He 75-80 %
- > Stars are composed mostly of He, are more massive and evolve faster, e.g. $10M_{\odot}$ star with 75% He evolves ~ 10 times faster than normal star (24% He).
- \triangleright High He abundance increases Initial Mass Function: $\xi(M)' \sim 1.5 \times \xi(M)$
- > Star formation peaks at $z\sim10 \approx t_{lookback}\sim13.3$ Gyr: $SFR'(z)\sim2.3\times SFR(z)$
- \triangleright Number of stars: $N'(m) \sim 5 \times N(m)$
- \triangleright Number of black holes: $N'_{BH} \sim 10 \times N_{BH}$

For a review on mirror world see

Z. Berezhiani, 2005

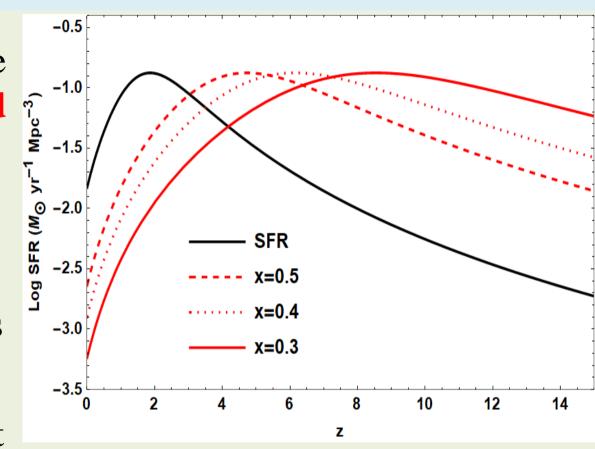
Star Formation Rate

In the period 0 < z < 14 more stars are formed in Mirror World relative to our:

$$\frac{\int_0^{14} SFR' \, dz}{\int_0^{14} SFR \, dz} = 2.3$$

The number of Mirror black holes $N'_{BH} \sim 10 N_{BH}$

Even if mirror matter does not make up all dark matter, or if formation of binary systems is not so efficient, the amplification factor still can be ~ 5.



SFR in the interval 0 < z < 14In red: Mirror SFR' for different temperatures x = T'/T.

(Madau & Dickinson 2014)

LIGO Signals from Mirror World

> Merging Rate with combining Mirror World amplification factors:

$$\mathcal{R}_{mirror} \sim 10 \times \mathcal{R}_{theor} \sim 50 - 100 \; \mathrm{Gp}c^{-3}yr^{-1}$$

coincides with **LIGO**'s bounds even if some assumptions of binary formation are relaxed;

- ➤ Hierarchical mergers are more probable in Mirror World and merger rates of upper mass gap systems (GW190521 & GW190426_190642) would agree better even with less strict assumptions;
- ➤ Production of 'heavy NSs' (GW190425) or lower mass gap objects (GW190814 & GW200210_092254) are easier in Mirror World, as they are dominated by He.

Conclusions:

In the Mirror World scenario:

- Number of binary systems is higher;
- BBH merger rate is amplified, coinciding better with LIGO estimations;
- Mass gap events could be better explained;
- Non-detection of EM-radiation is natural, since Mirror photons DO NOT interact with Ordinary particles.

Prediction: Binary compact objects' merger rates are order of 10 higher than expected and only 1 of 10 NS-NS events discovered by GW detectors may have EM-counterpart.

References:

- 1. "Gravitational Waves from Mirror World", MDPI Physics 1 (2019) 67;
- 2. "LIGO Signals from the Mirror World", MNRAS 487 (2019) 650;
- 3. "Binary Neutron Star Mergers with Missing Electromagnetic Counterparts as Manifestations of Mirror World", Phys. Lett. B 804 (2020) 135402;
- 4. "Unexpected LIGO events and the Mirror World", MNRAS 503 (2021) 2882.

