

# Solid state or Semiconductor Detectors

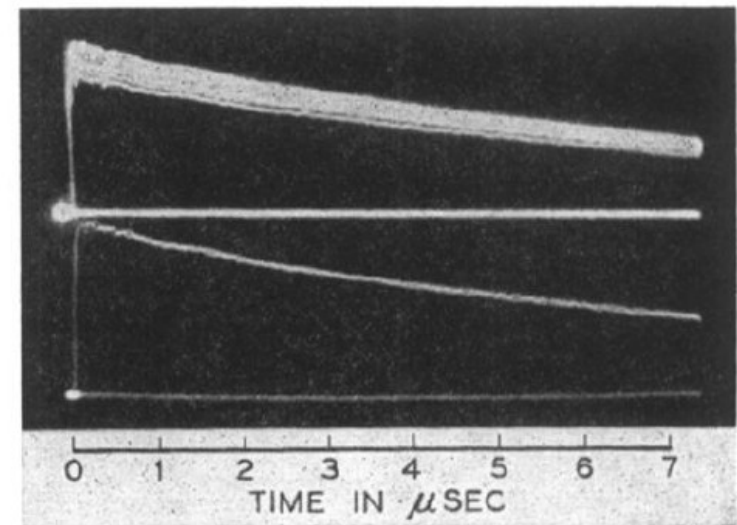
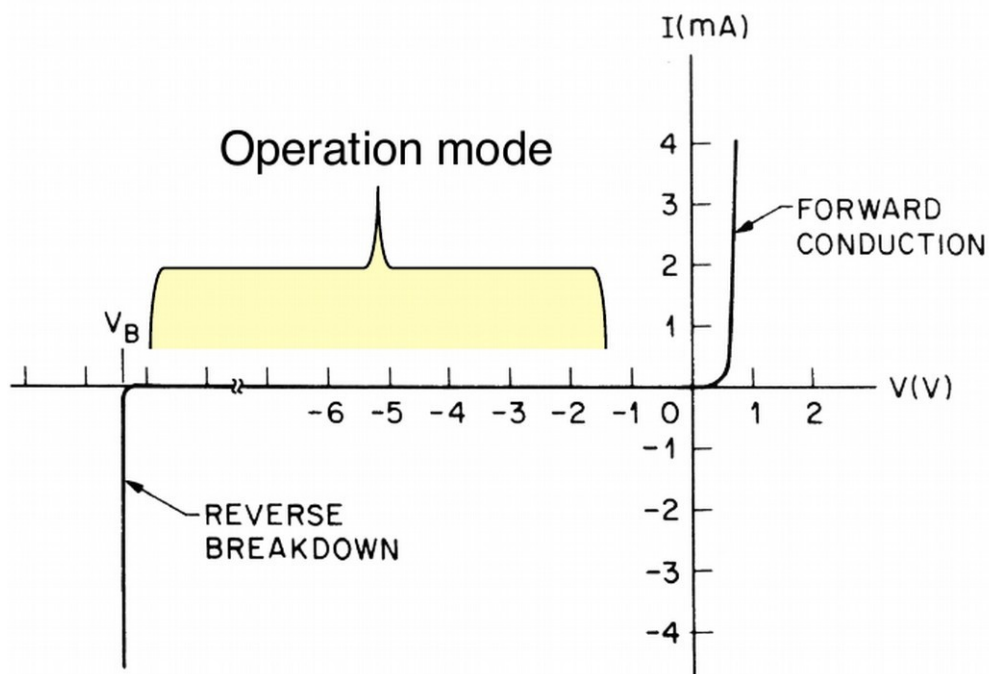
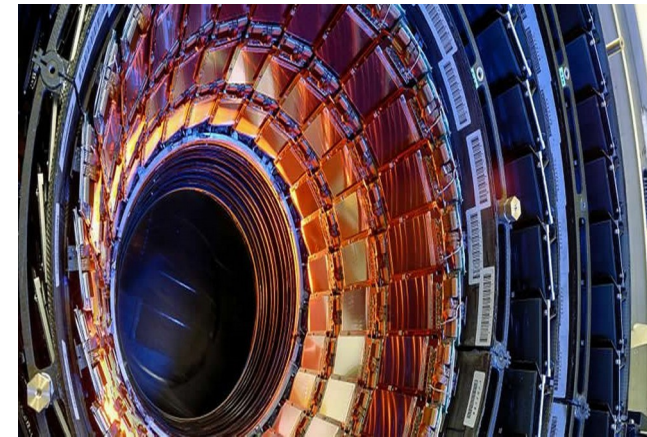
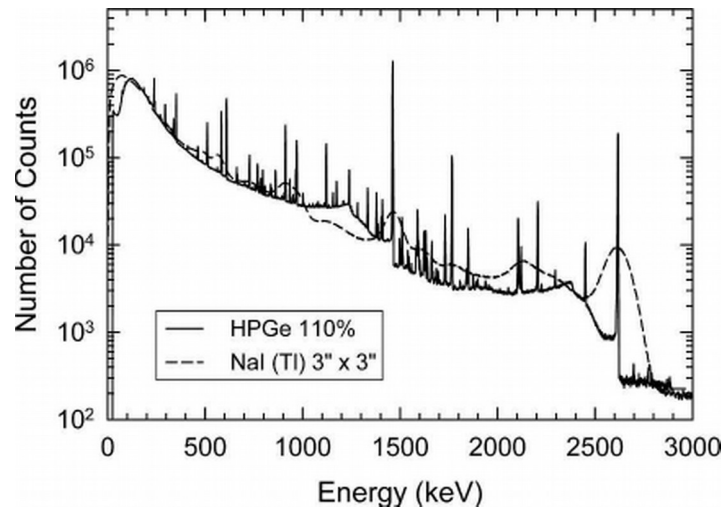
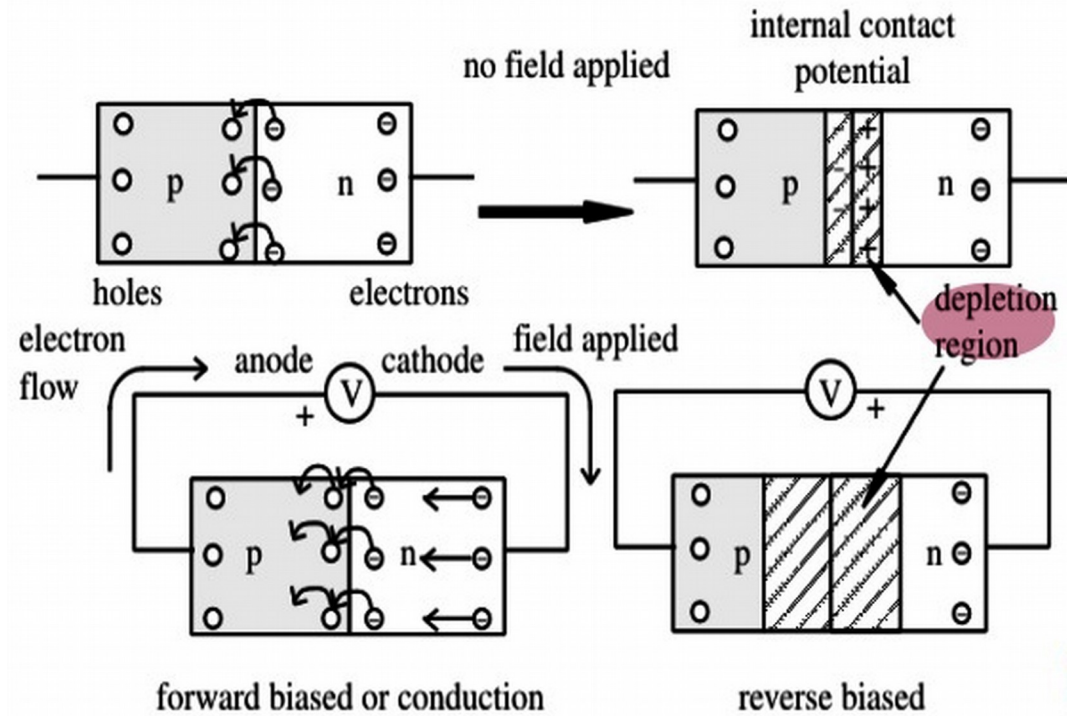
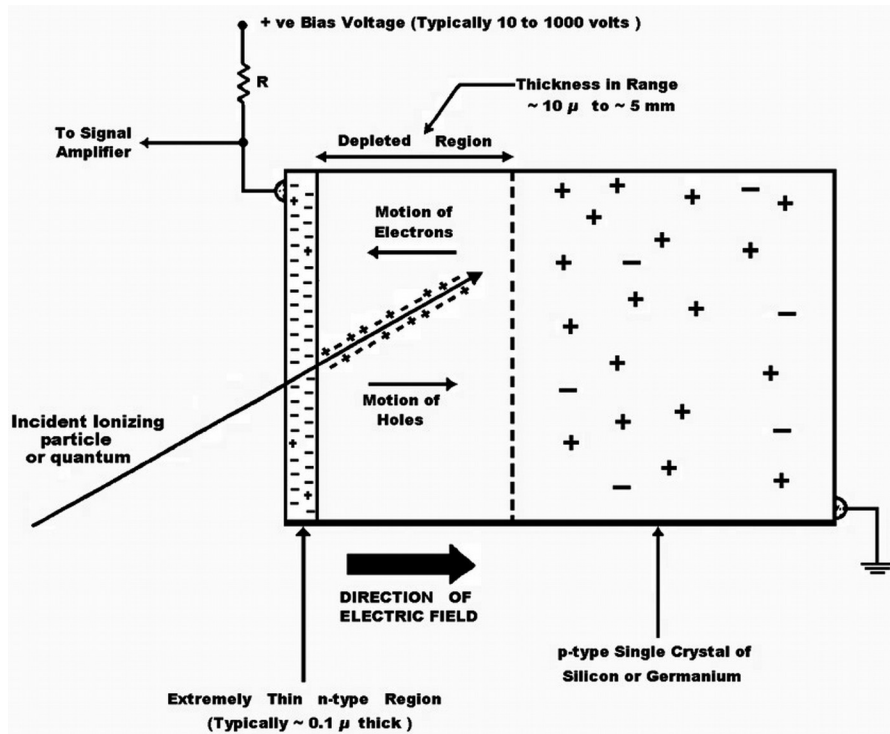
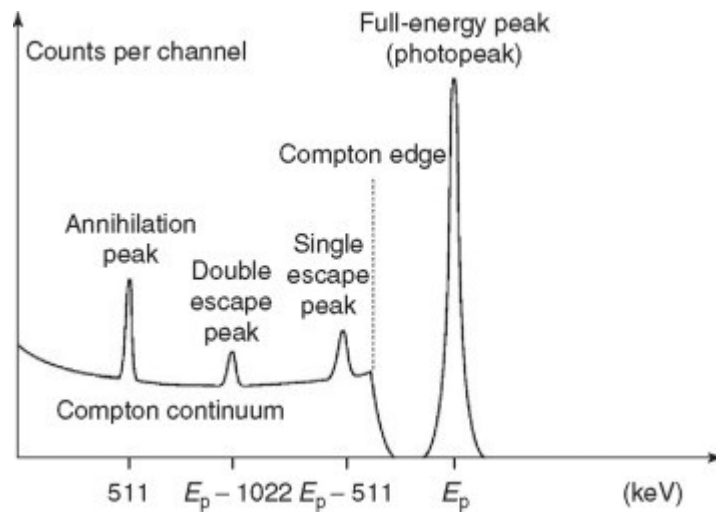
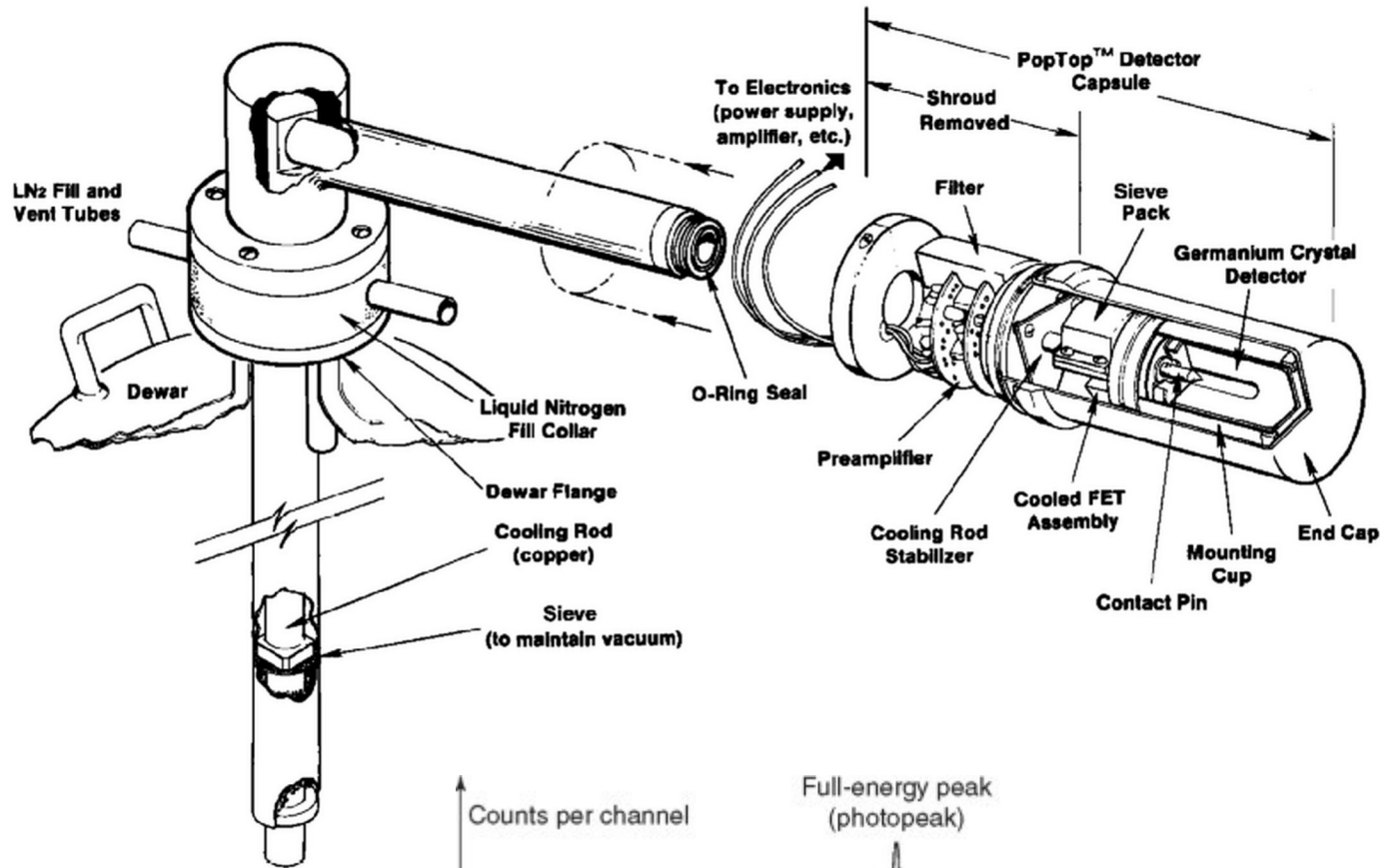


FIG. 3. Photograph of pulses from sixteen alpha-particles striking the  $n-p$  barrier.

# Semiconductor detectors



# SemiCons for Gamma Spectroscopy

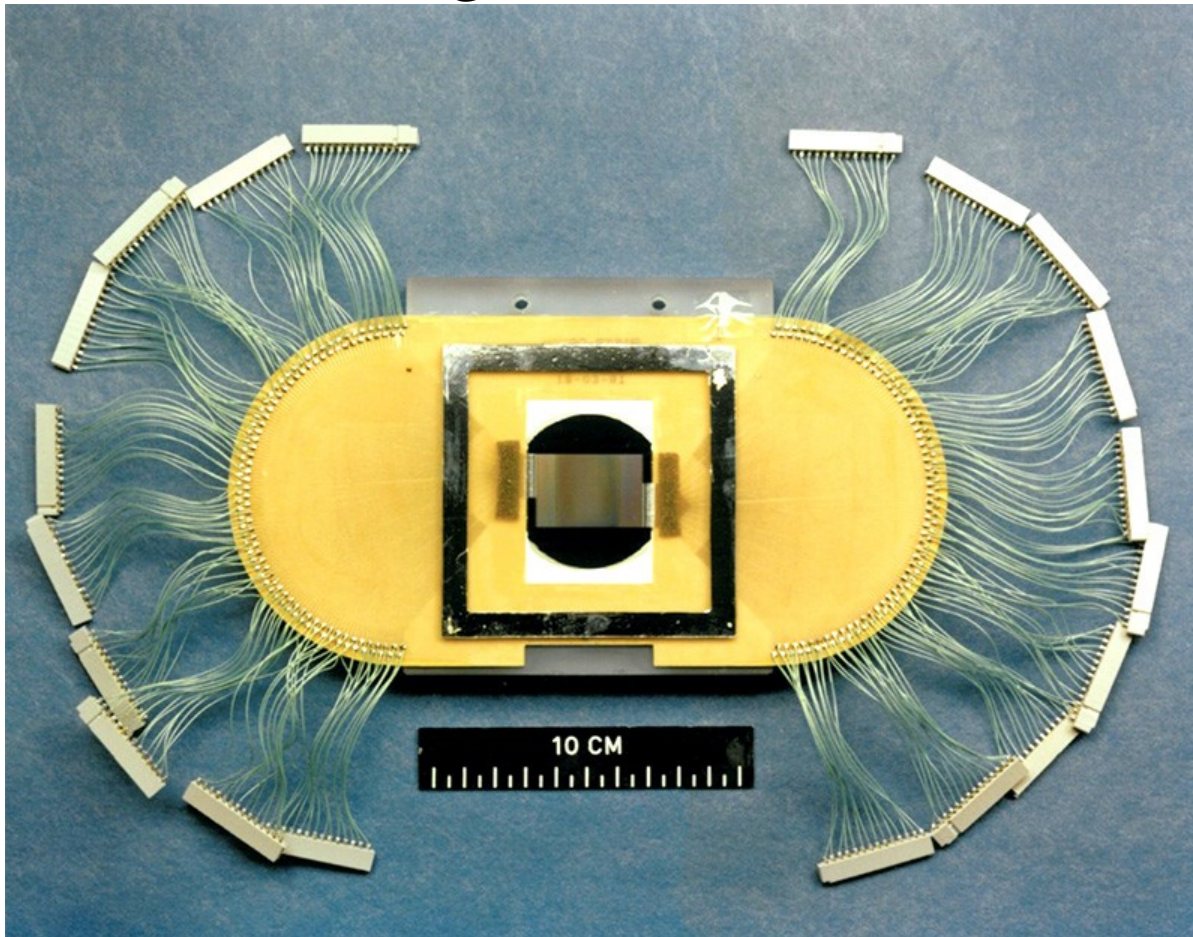




# First Silicon Strip Detector

Goal: Measure lifetime and mass of the charm mesons  $D^0$ ,  $D^-$ ,  $D^+$ ,  $D_s^+$ ,  $D_s^-$

The NA11 / NA32 @ CERN 1983



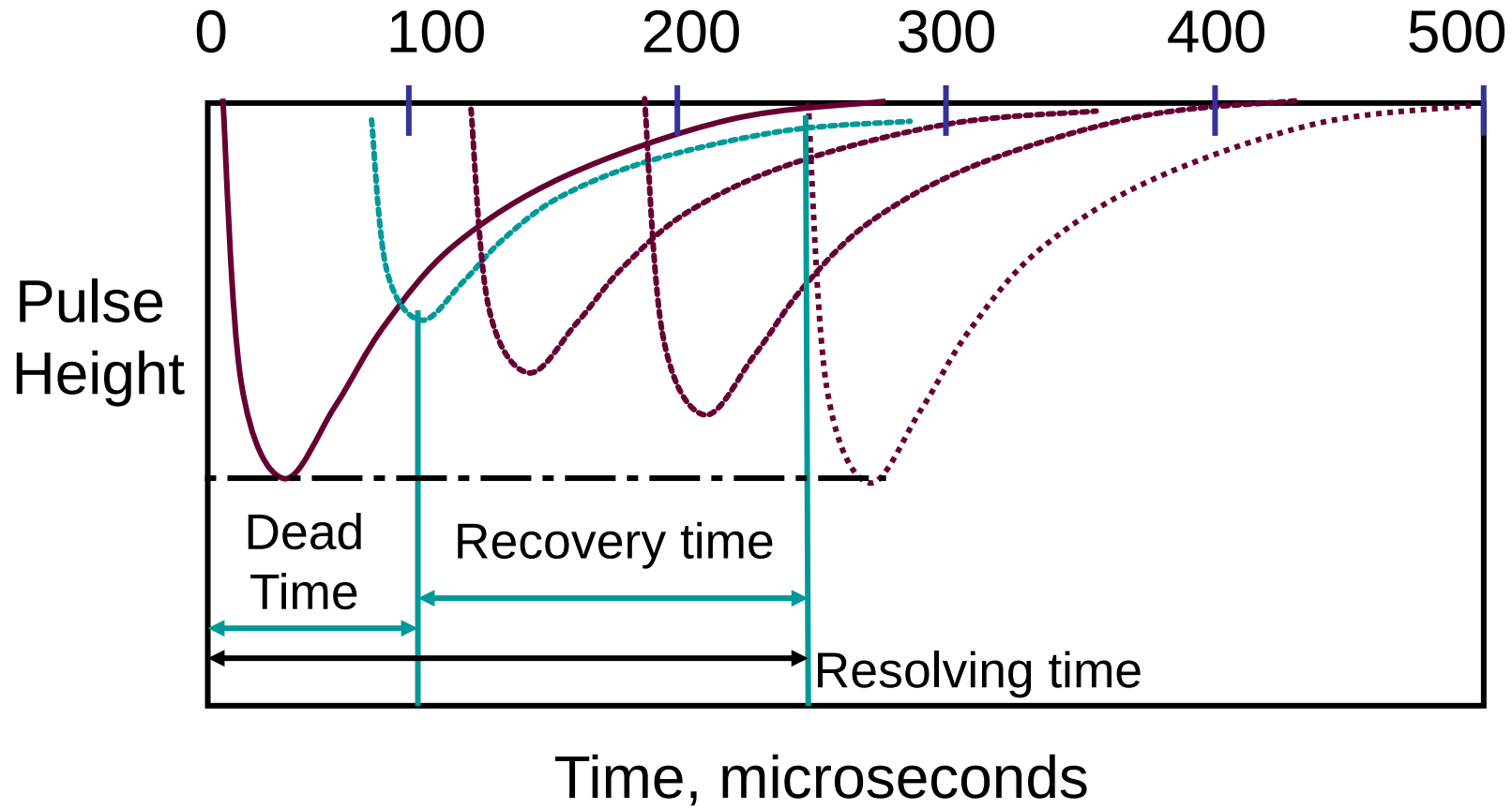
NIM205 (1983) 99

Surface 24 cm<sup>2</sup> (2" wafer)  
1200 strip, 20  $\mu$ m pitch  
Ever 3rd/6th strip connected.  
**Precision 4,5  $\mu$ m**

8 silicon detectors!  
2 in front,  
6 behind the target

Ratio detector surface !  
to nearby electronics!  
**surface 1:300**





# Detector Dead Time Correction

For random hit rate &

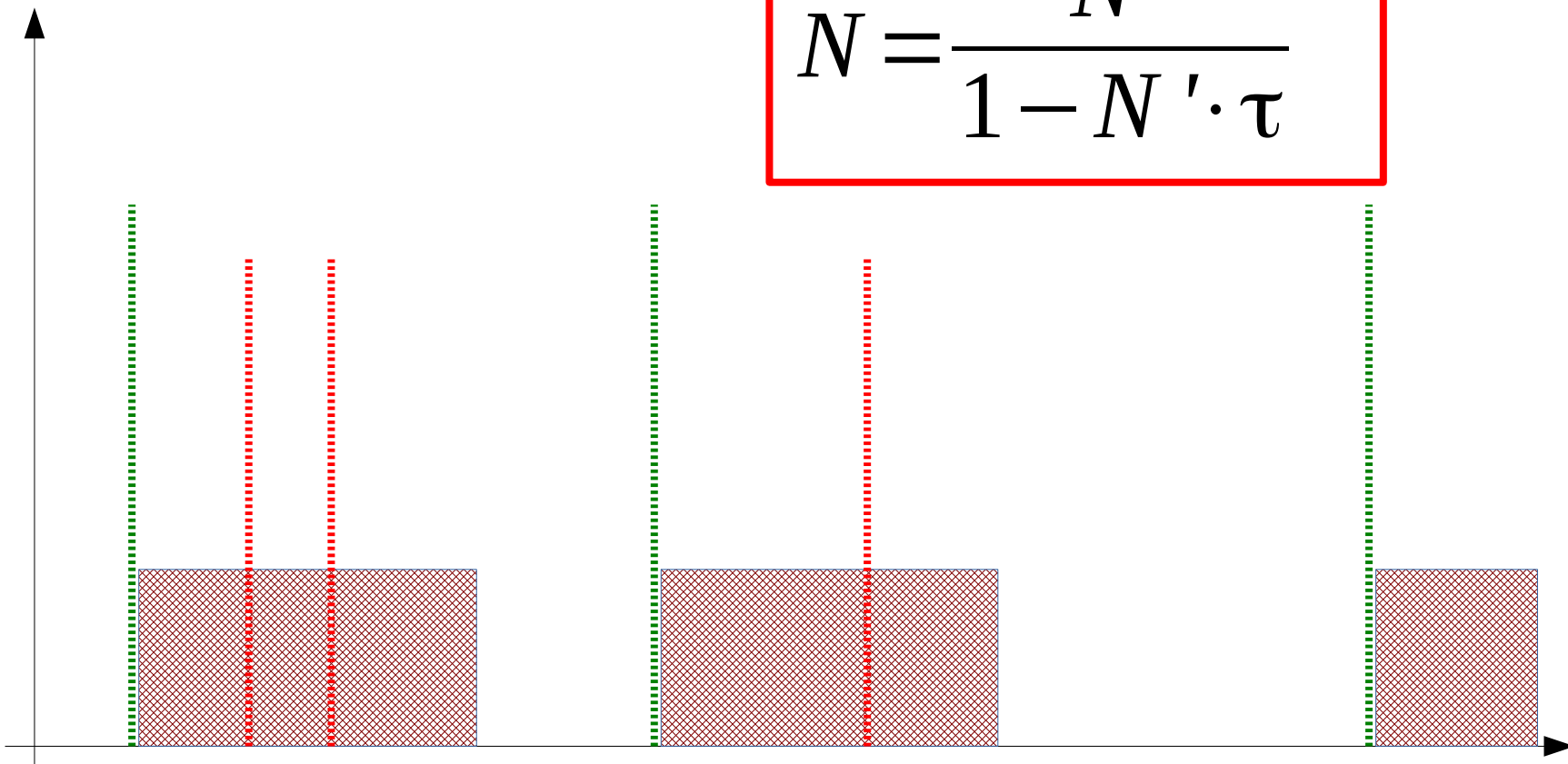
Constant  $\tau$

$N$  = real number of events

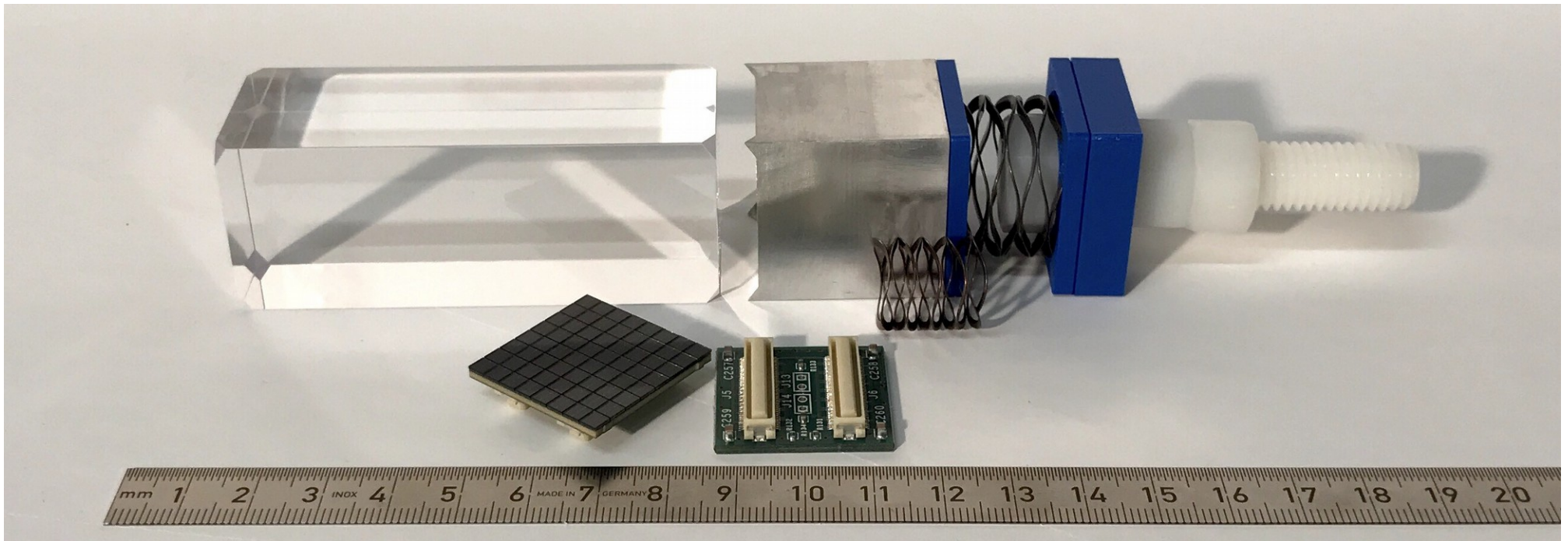
$N'$  = Recorded number of events

$$\frac{N}{N'} = \frac{1}{1 - N' \cdot \tau}$$

$$N = \frac{N'}{1 - N' \cdot \tau}$$



# Scintillating Detectors





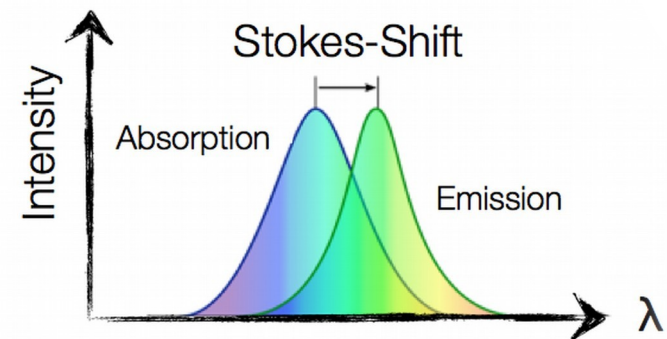
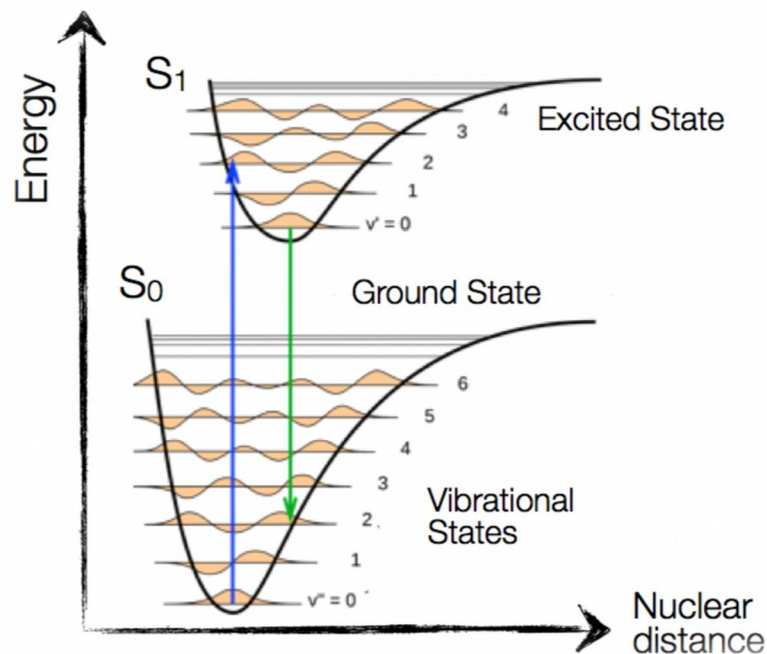
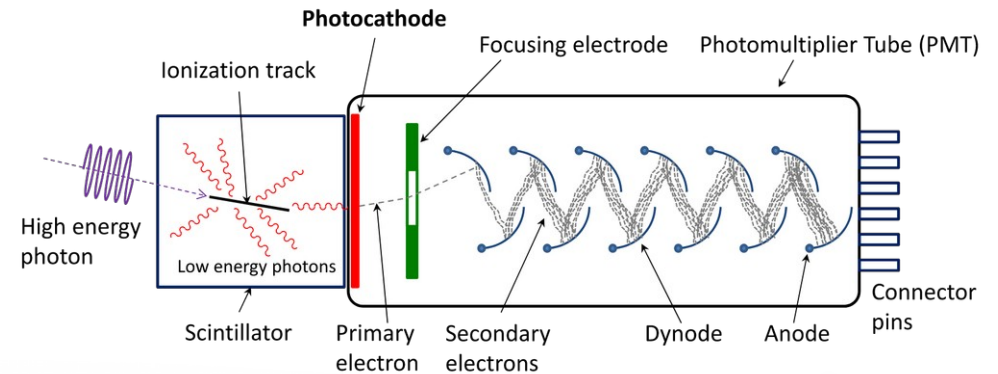
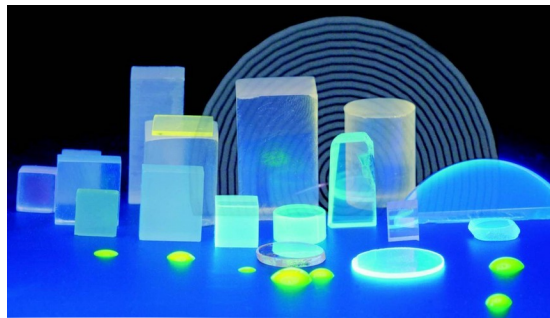
# Scintillation detectors

The light emission is governed by the electronic transitions in the molecule.

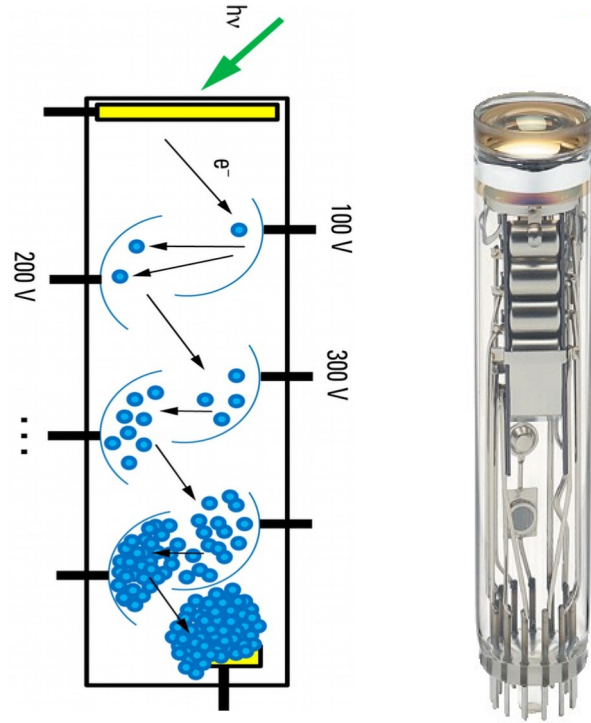
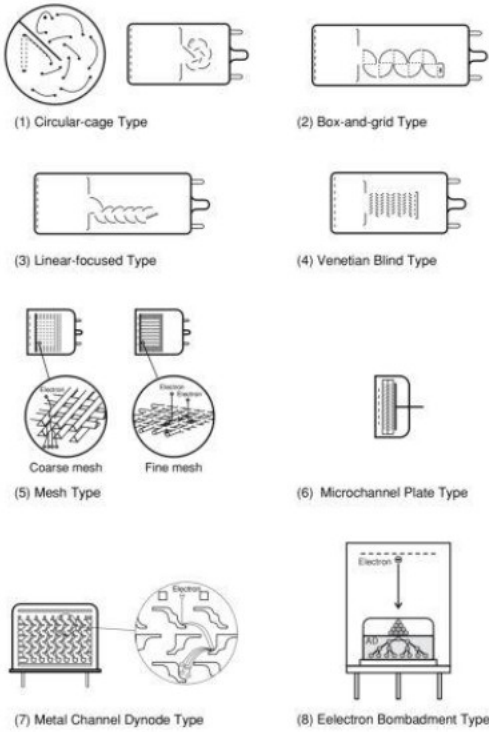
The electronic levels have a typical energy spacing of  $\sim 4$  eV.

The vibrational levels of the molecule ( $dE \sim 0.2$  eV) also play a role.

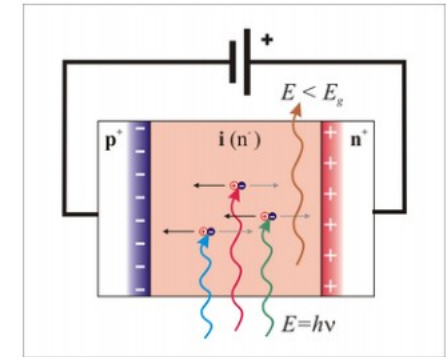
Electrons in high levels typically deexcite to the lowest excited state without emission of radiation.



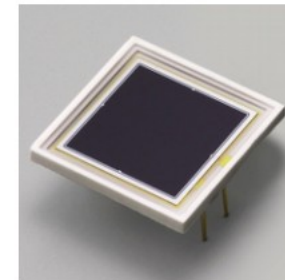
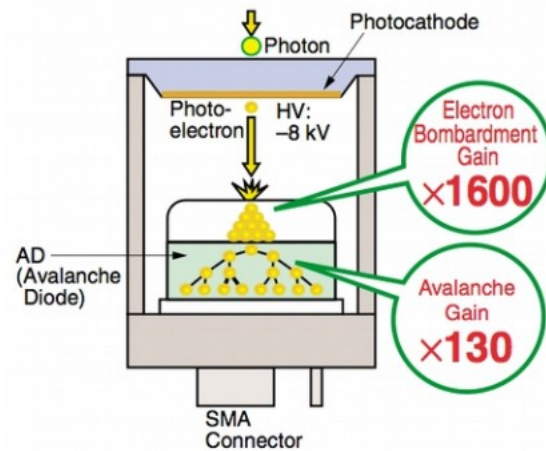
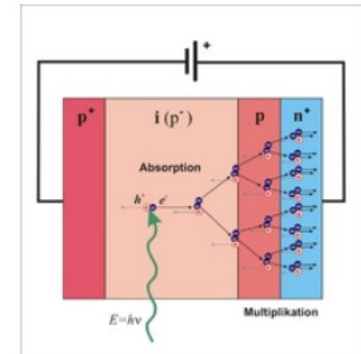
# Photosensitive sensors



PIN

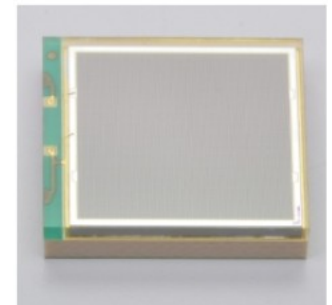


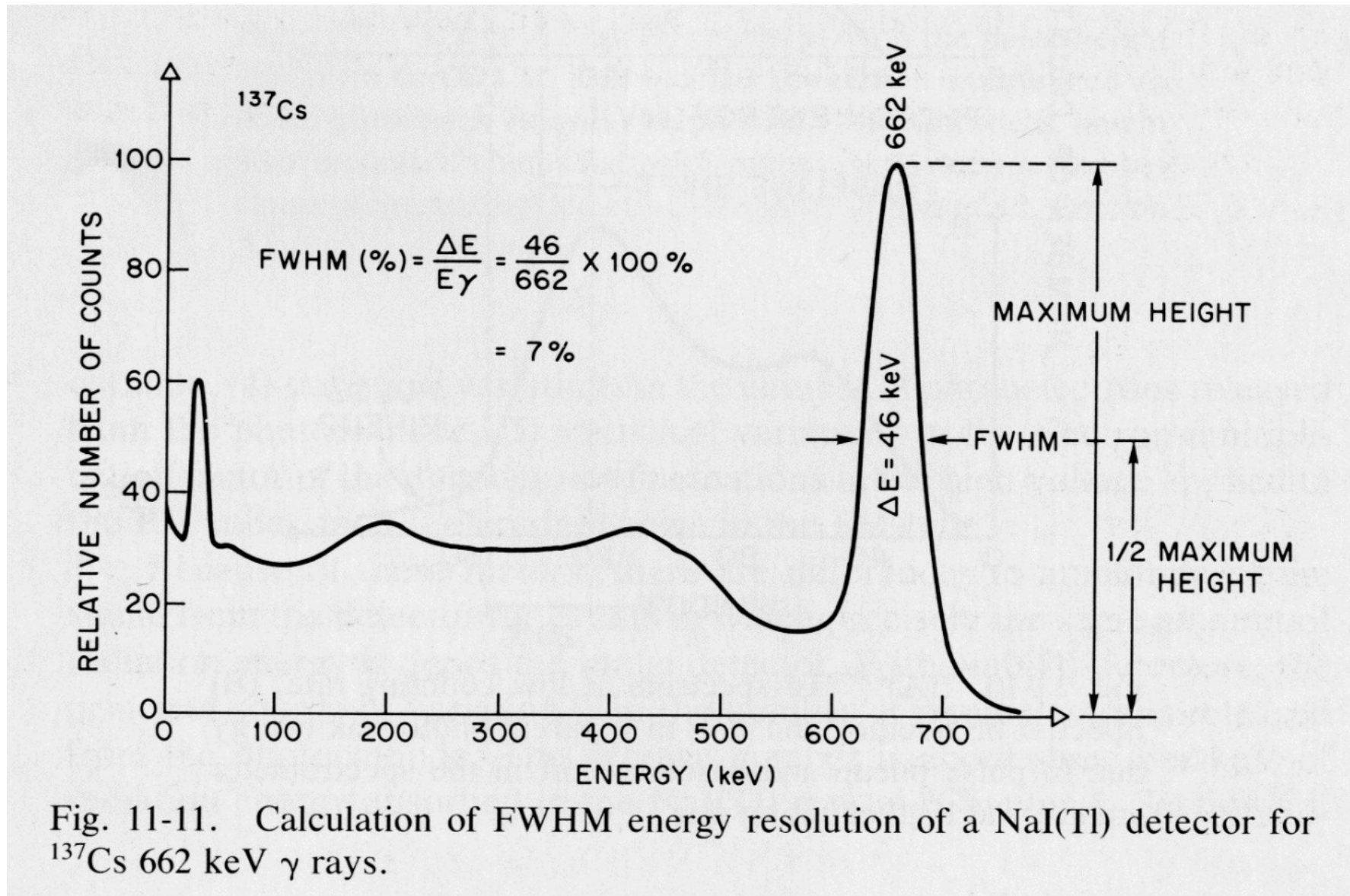
APD



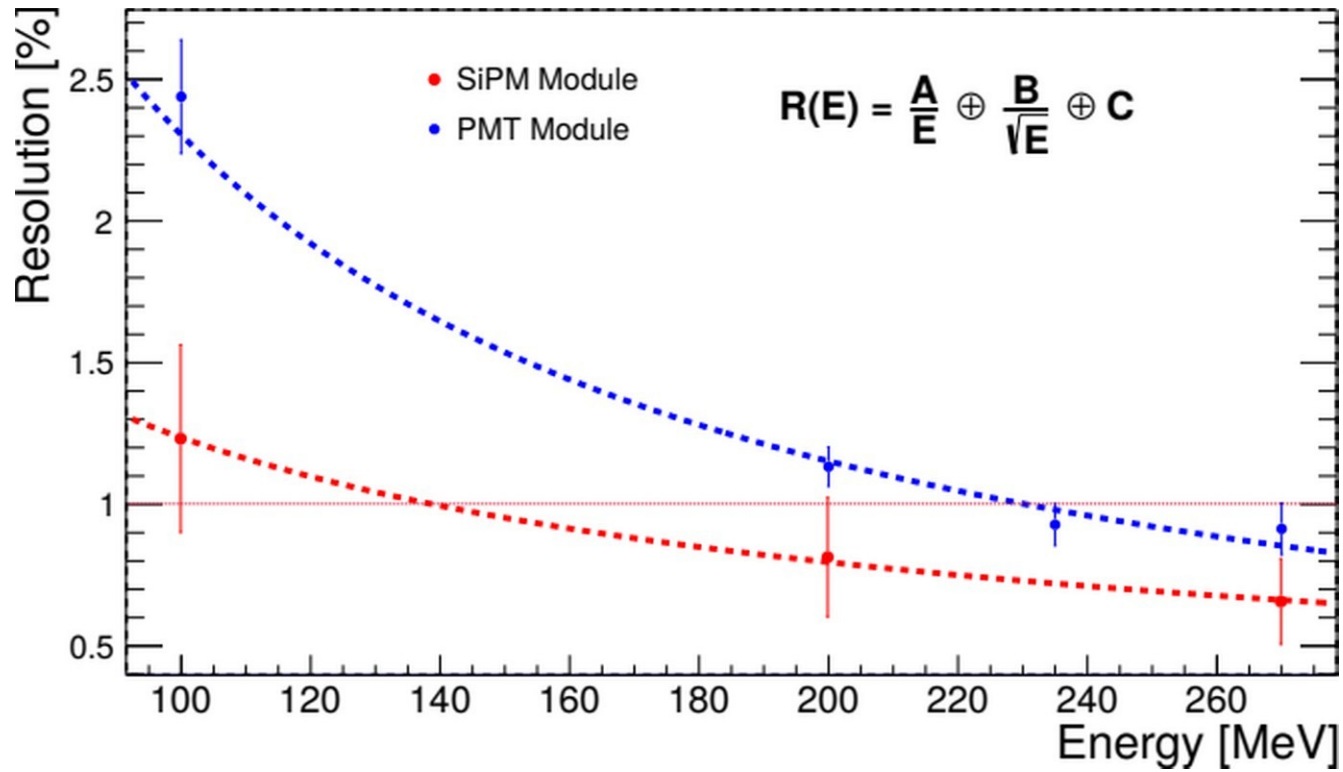
PIN/APD

SiPM



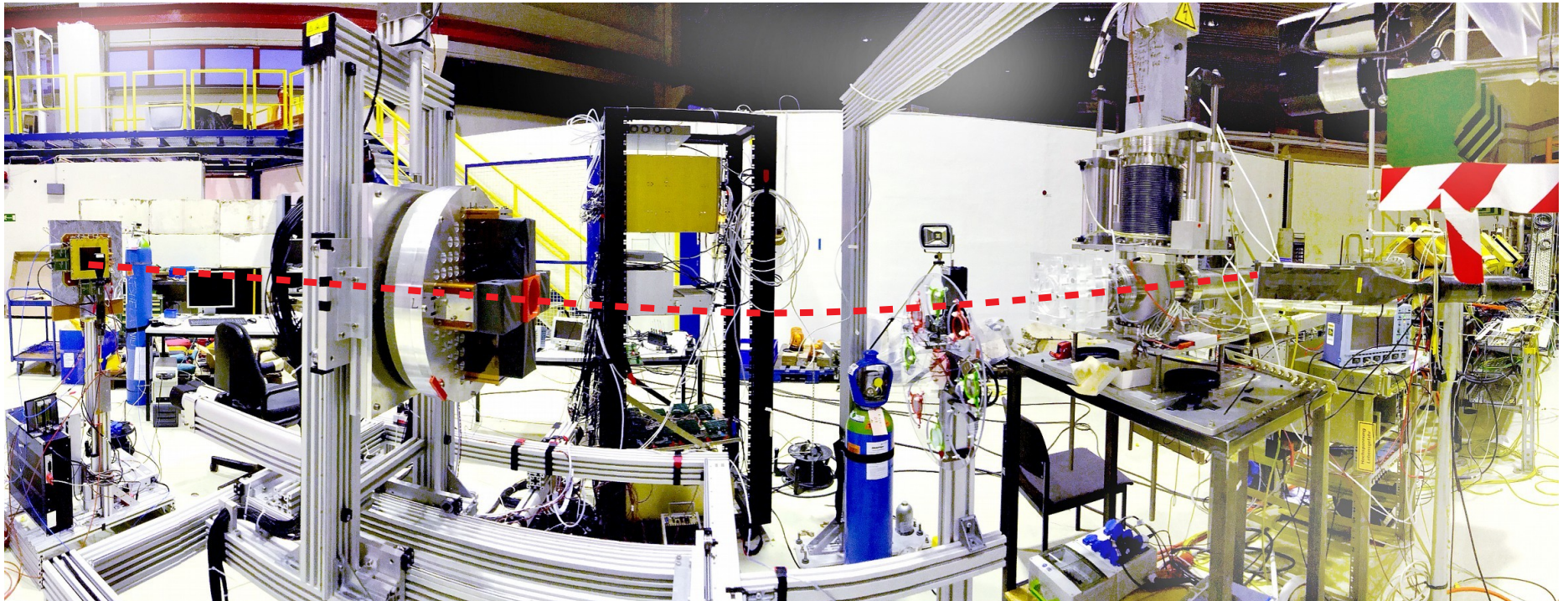
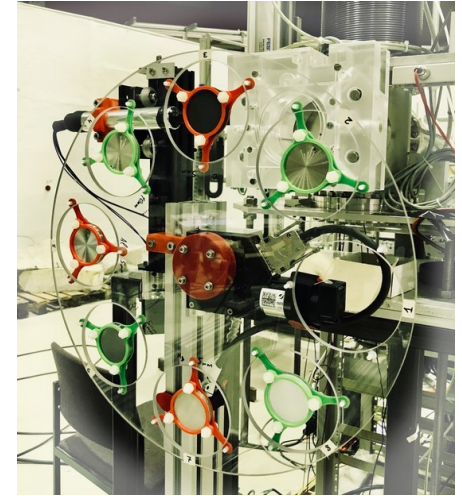
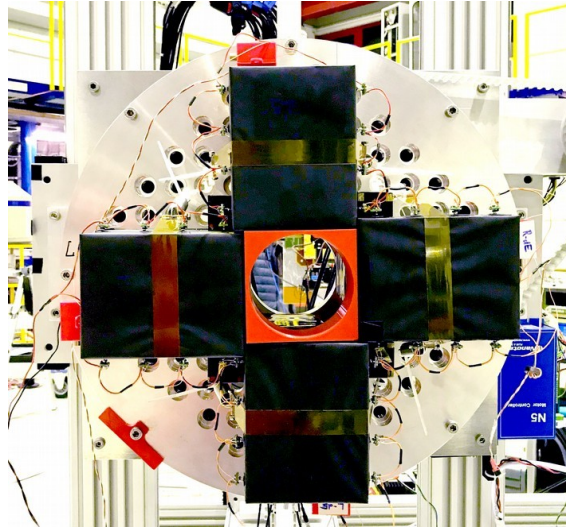
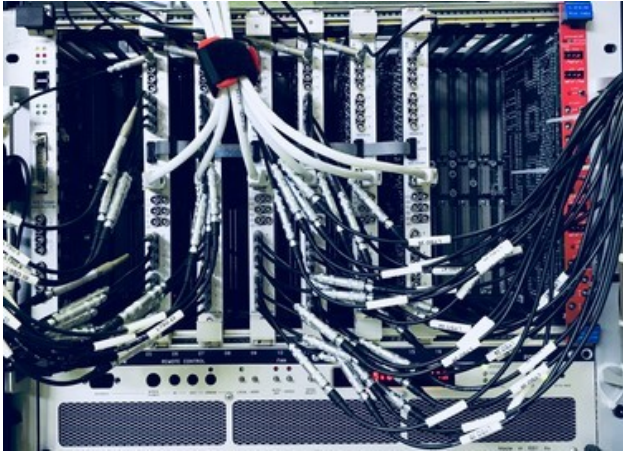








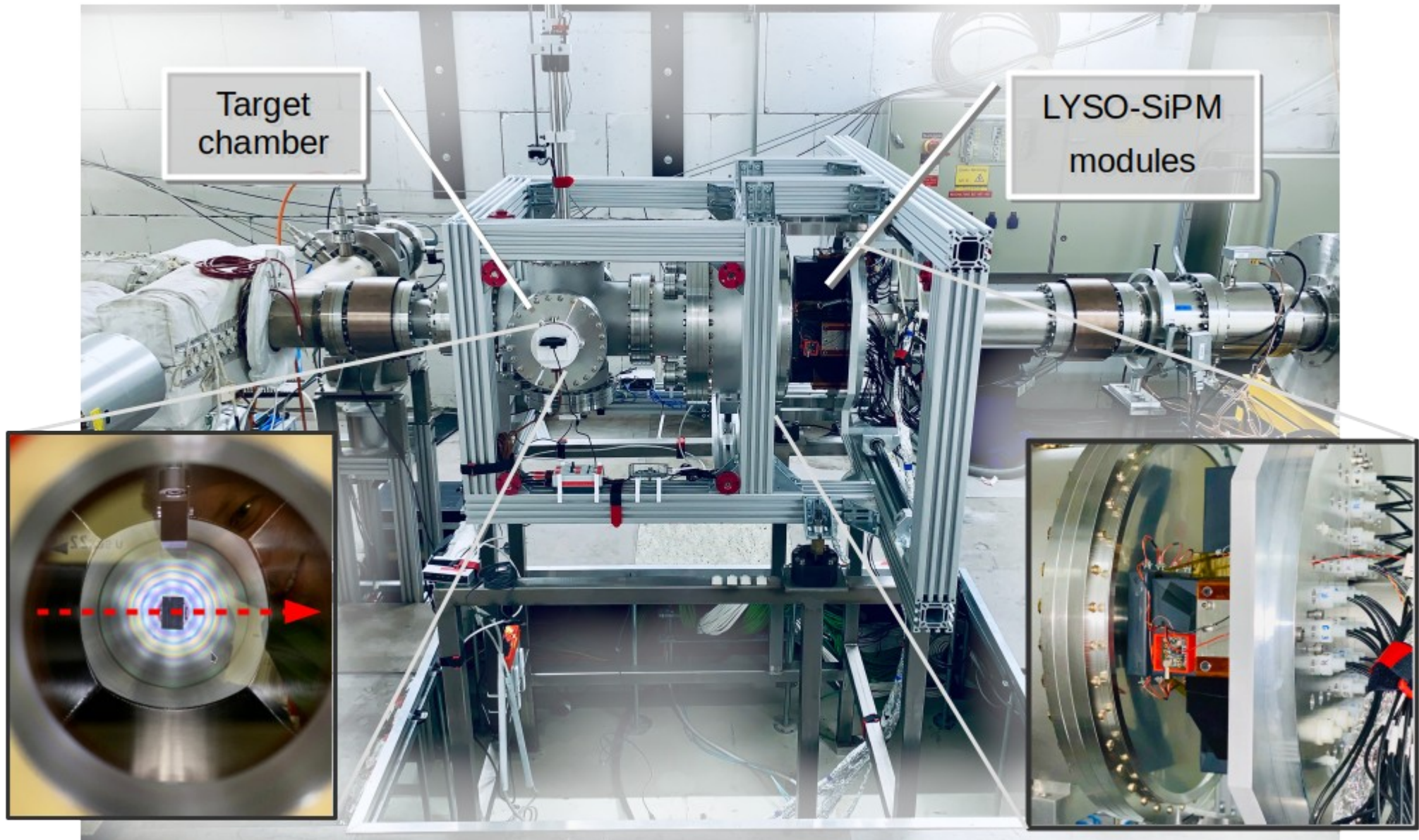
# Jedi Polarimeter



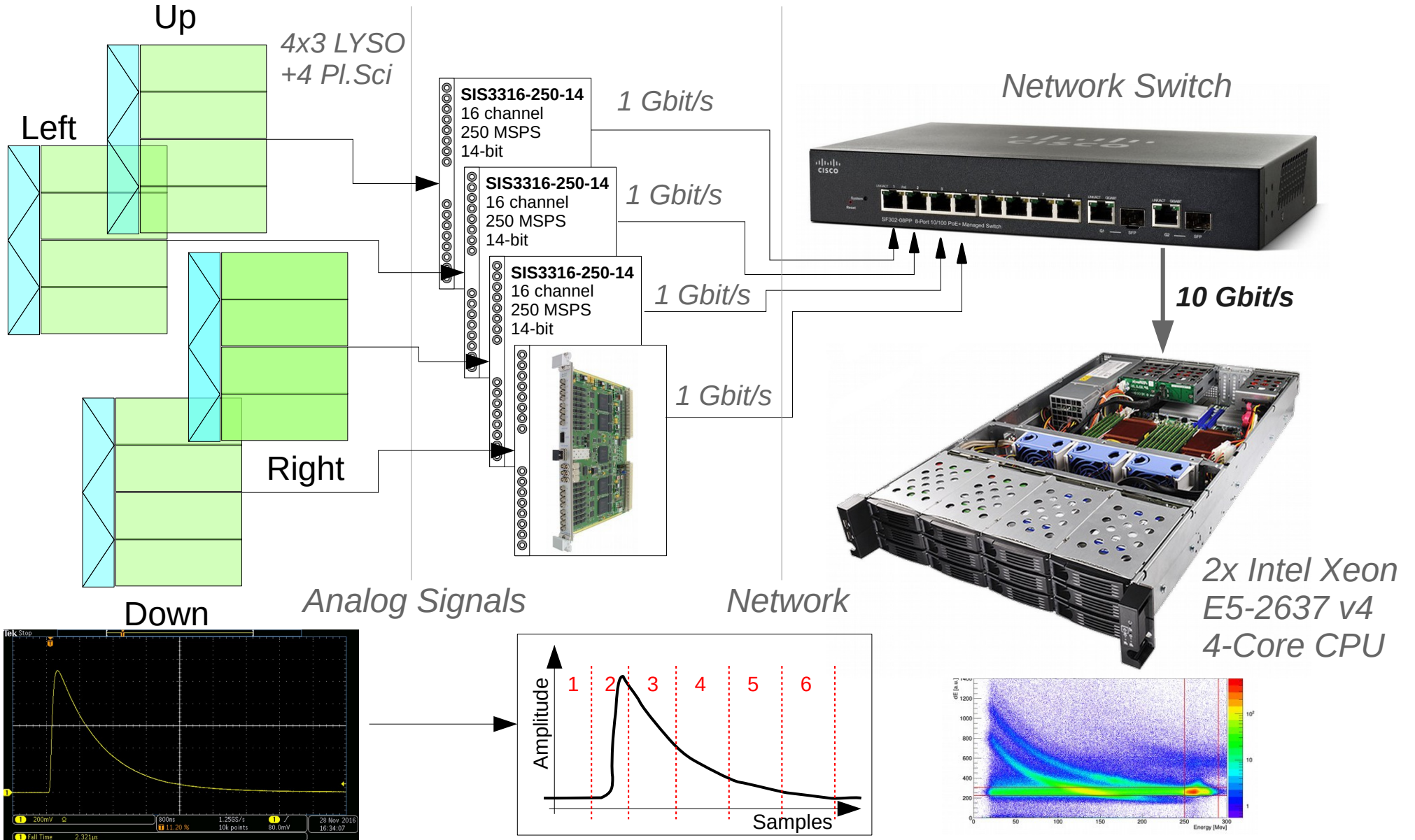








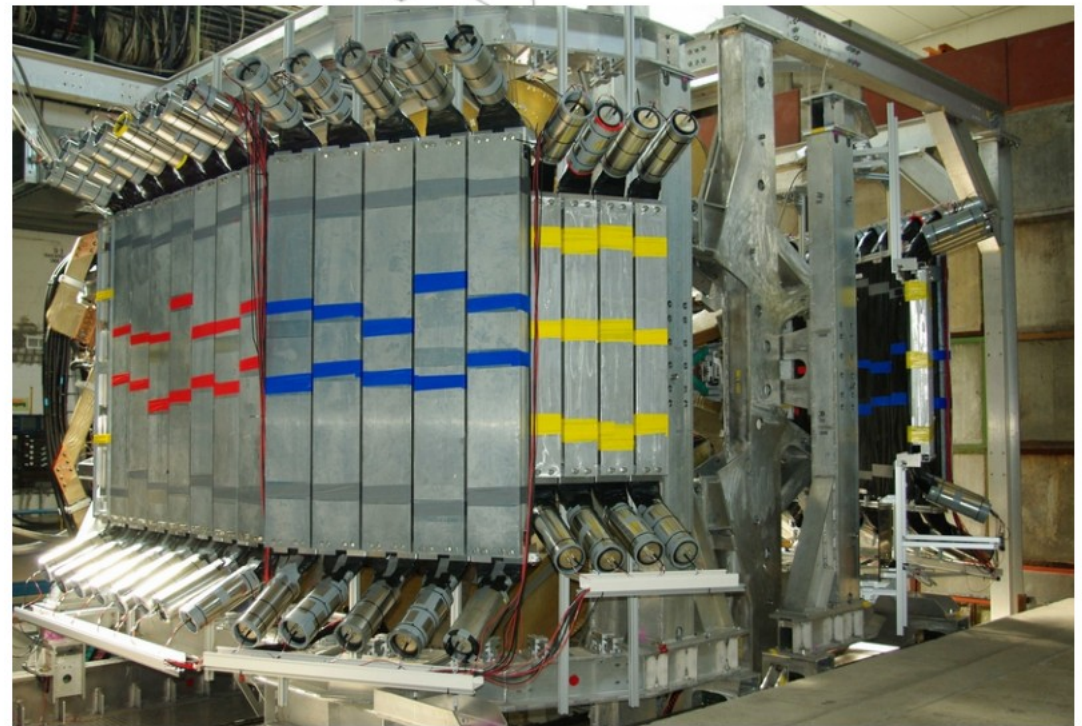
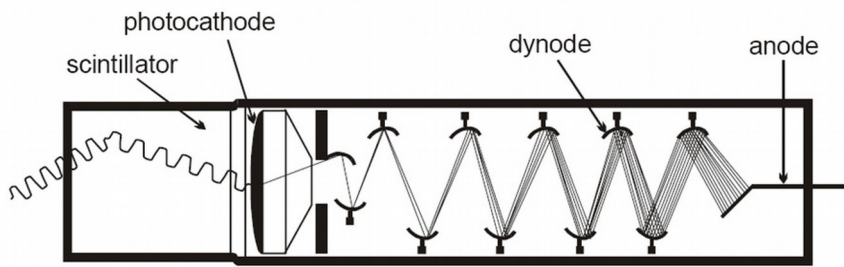
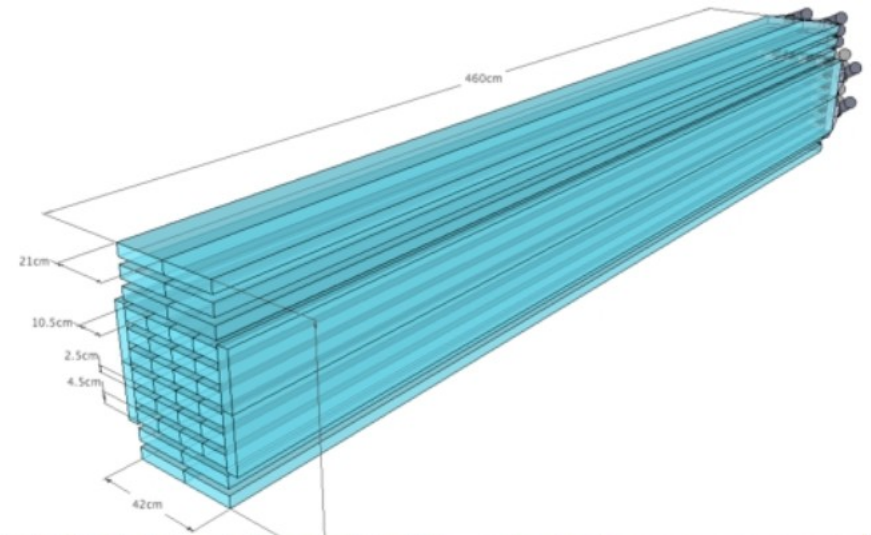
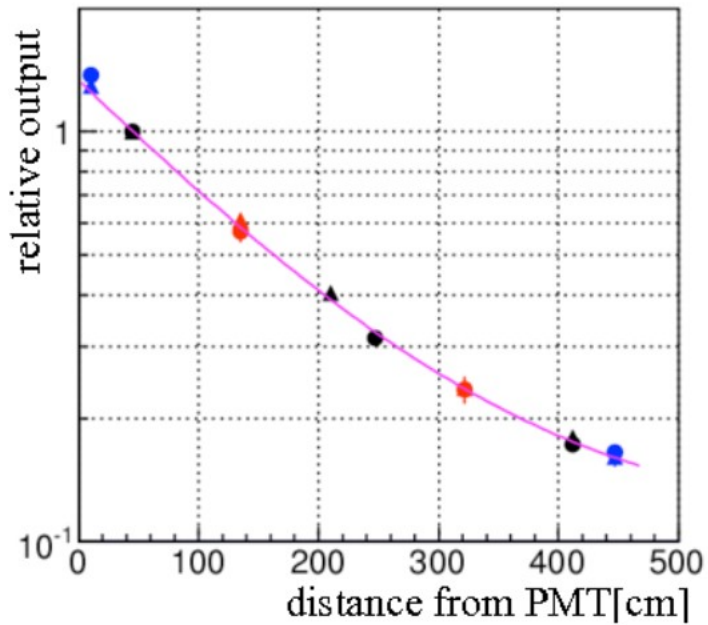
# Readout of JePo





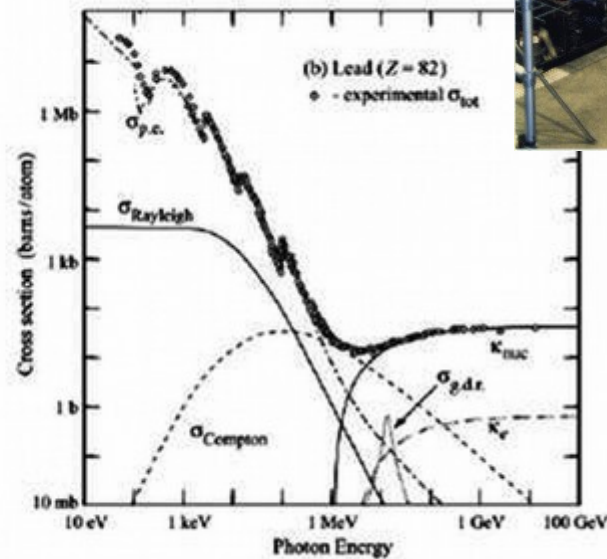
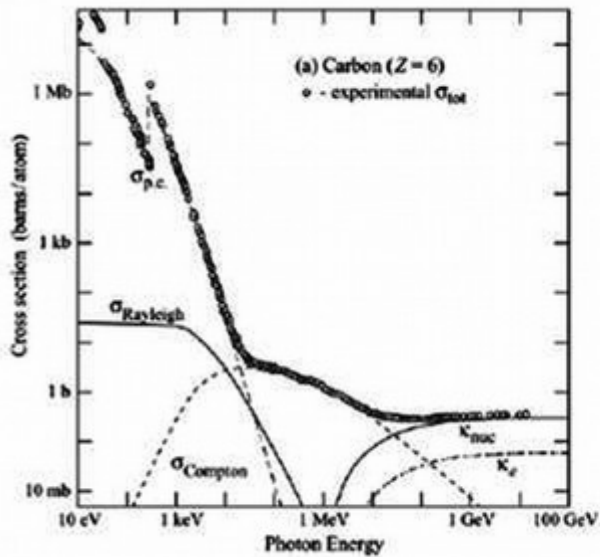
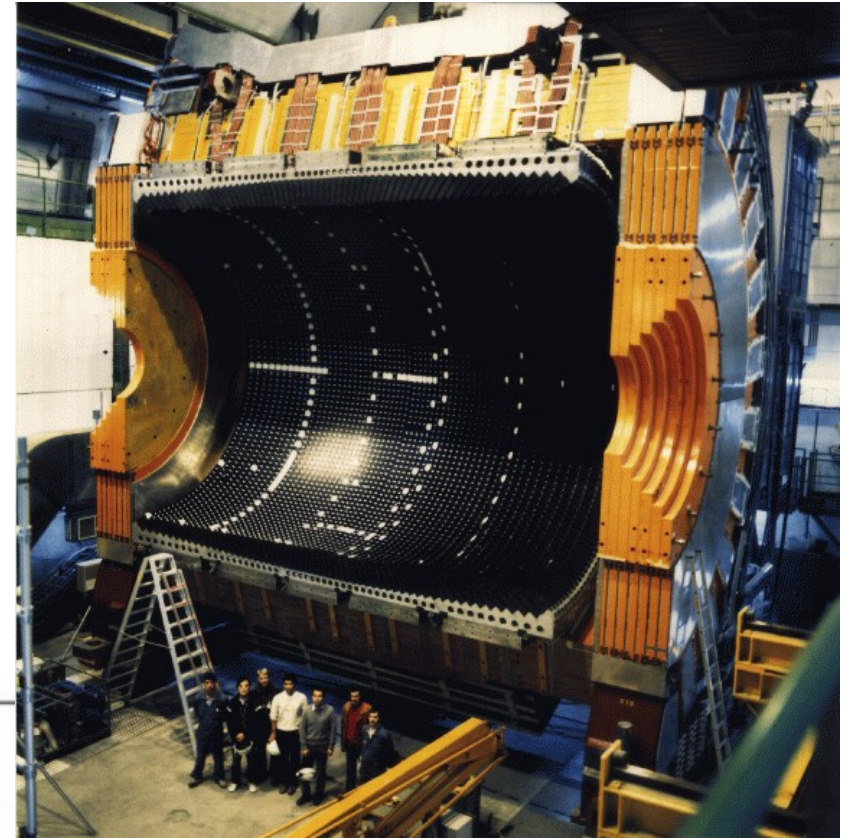
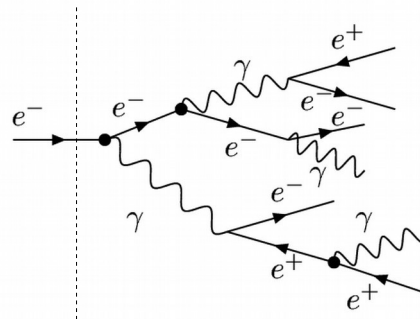
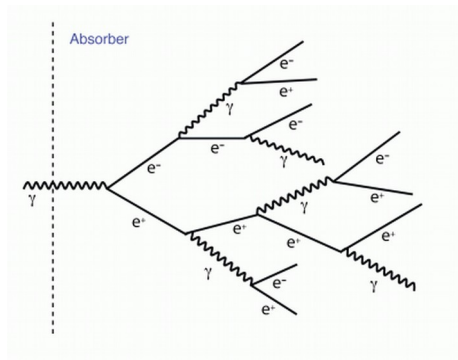
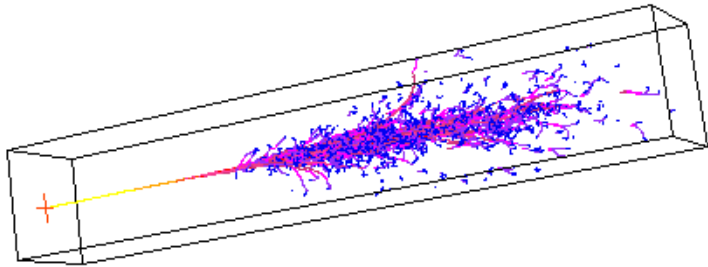
# TOF—time of flight technique

$$Y = Y_0 \cdot e^{\frac{-l}{\lambda}}$$

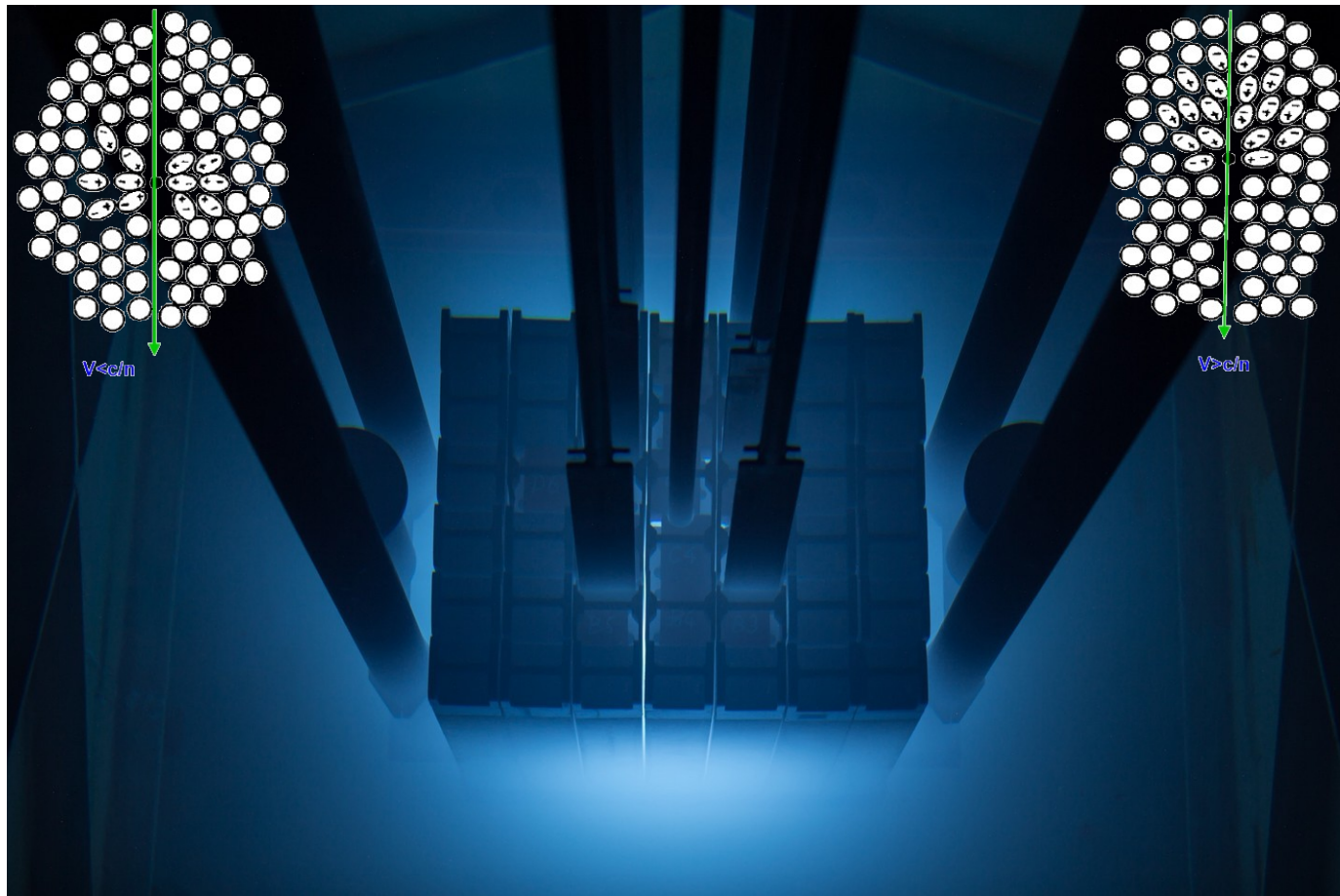




# EMC—electro magnetic calorimeters



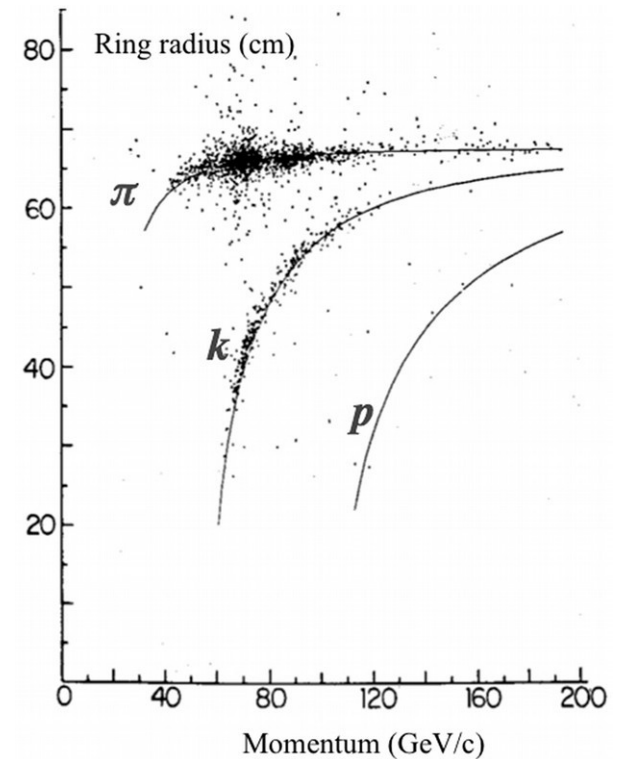
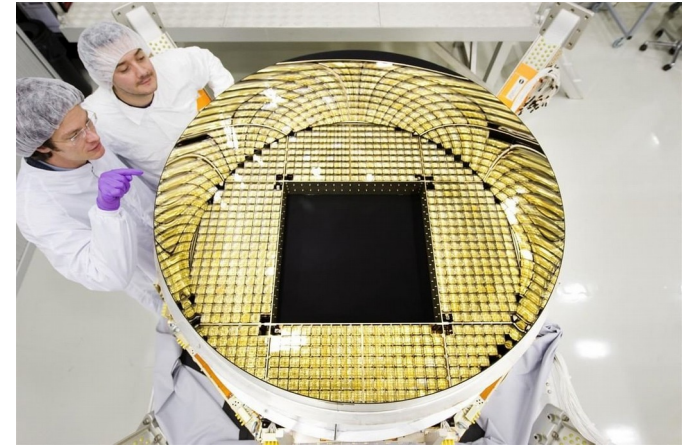
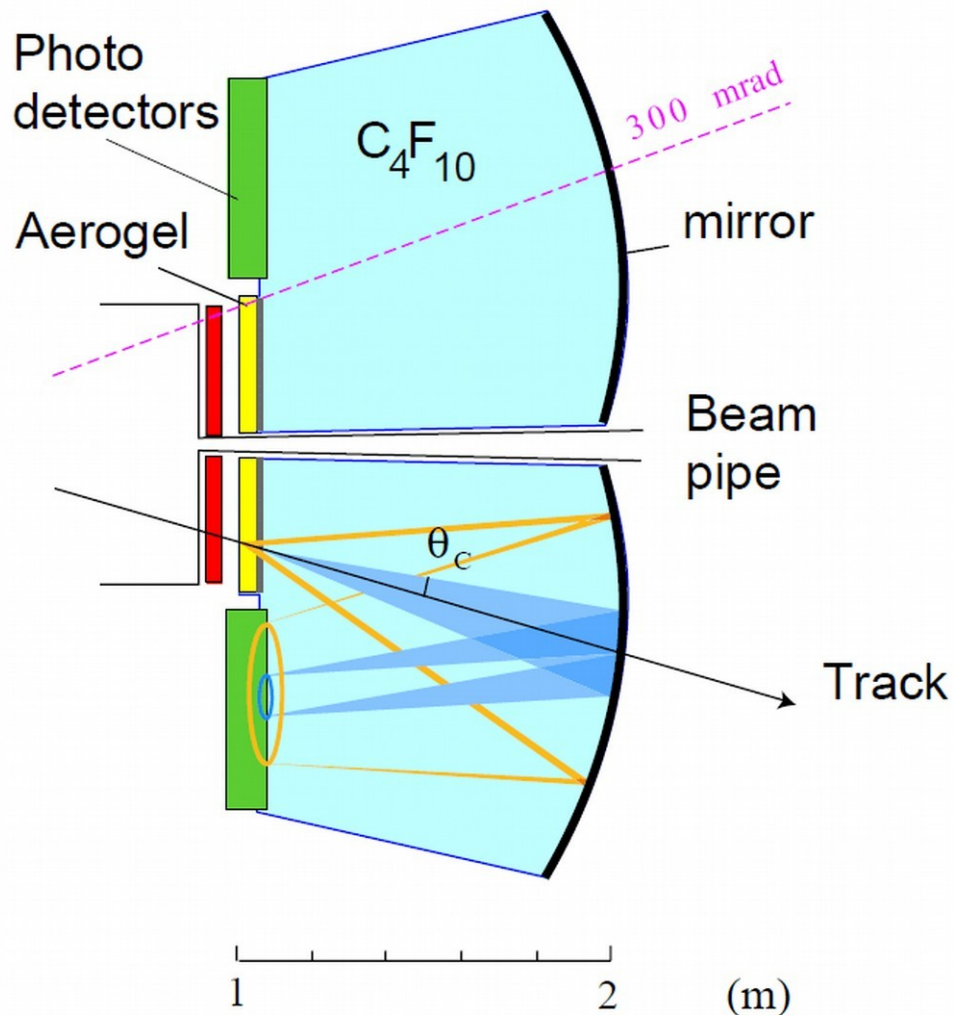
# Cherenkov Detectors





# RICH – Ring Image Cherenkov

$$N \propto 1 - \frac{1}{n^2 \beta^2} = 1 - \frac{1}{n^2} \left( 1 + \frac{m^2}{p^2} \right)$$

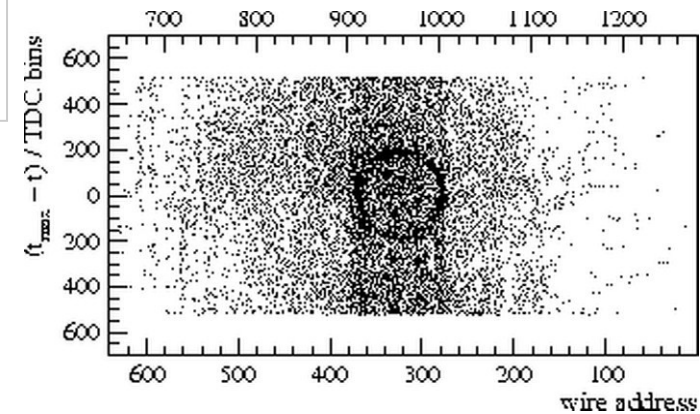
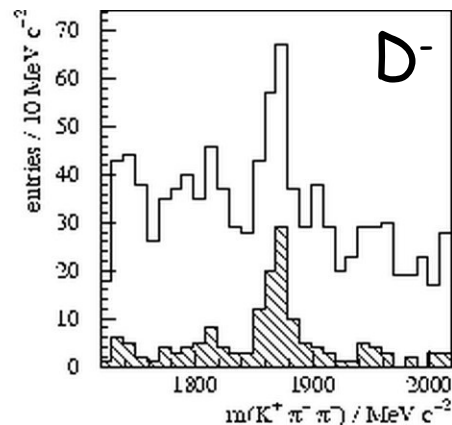
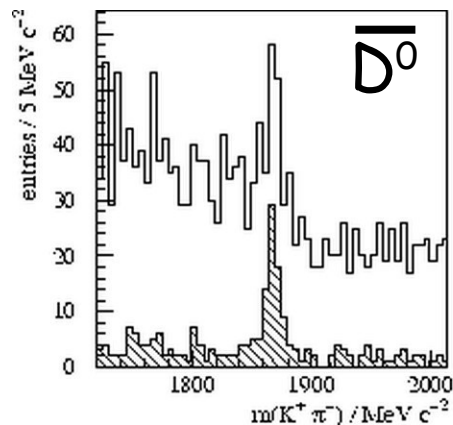
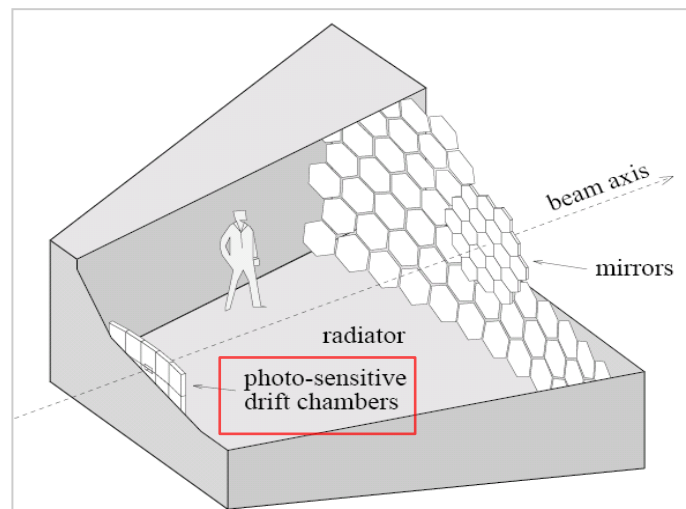
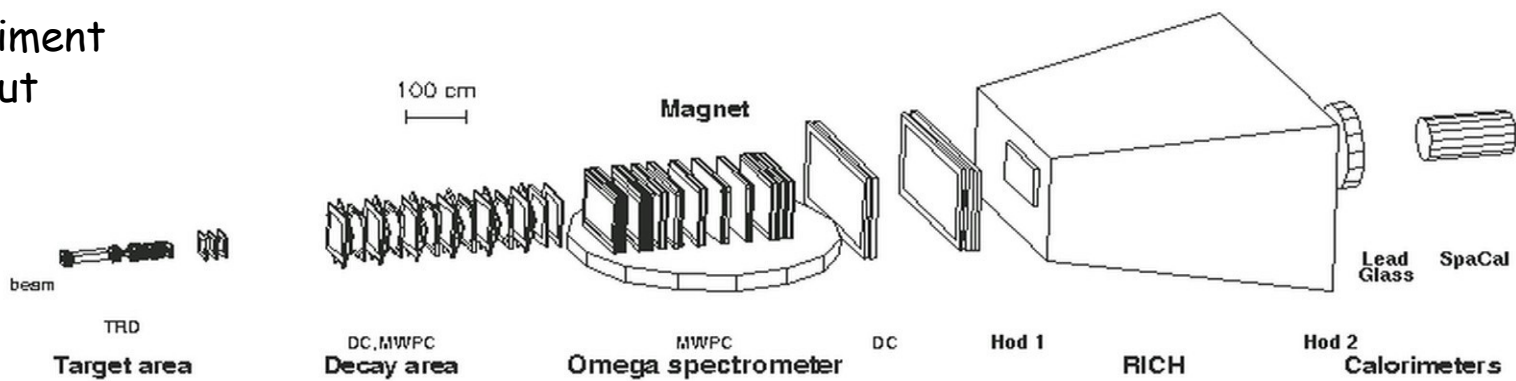




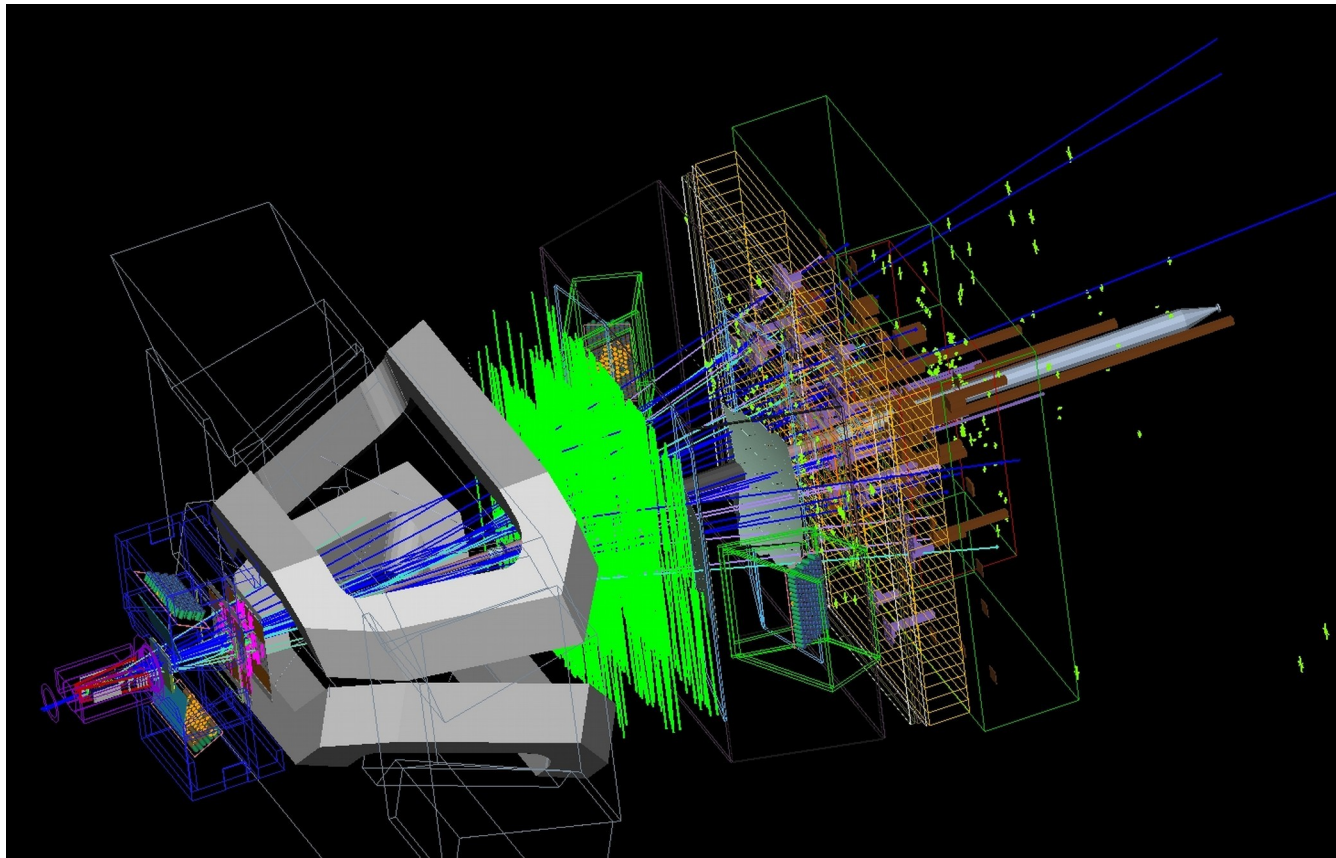
# RICH -- Ring Image Cherenkov detectors

WA89 experiment  
1993 layout

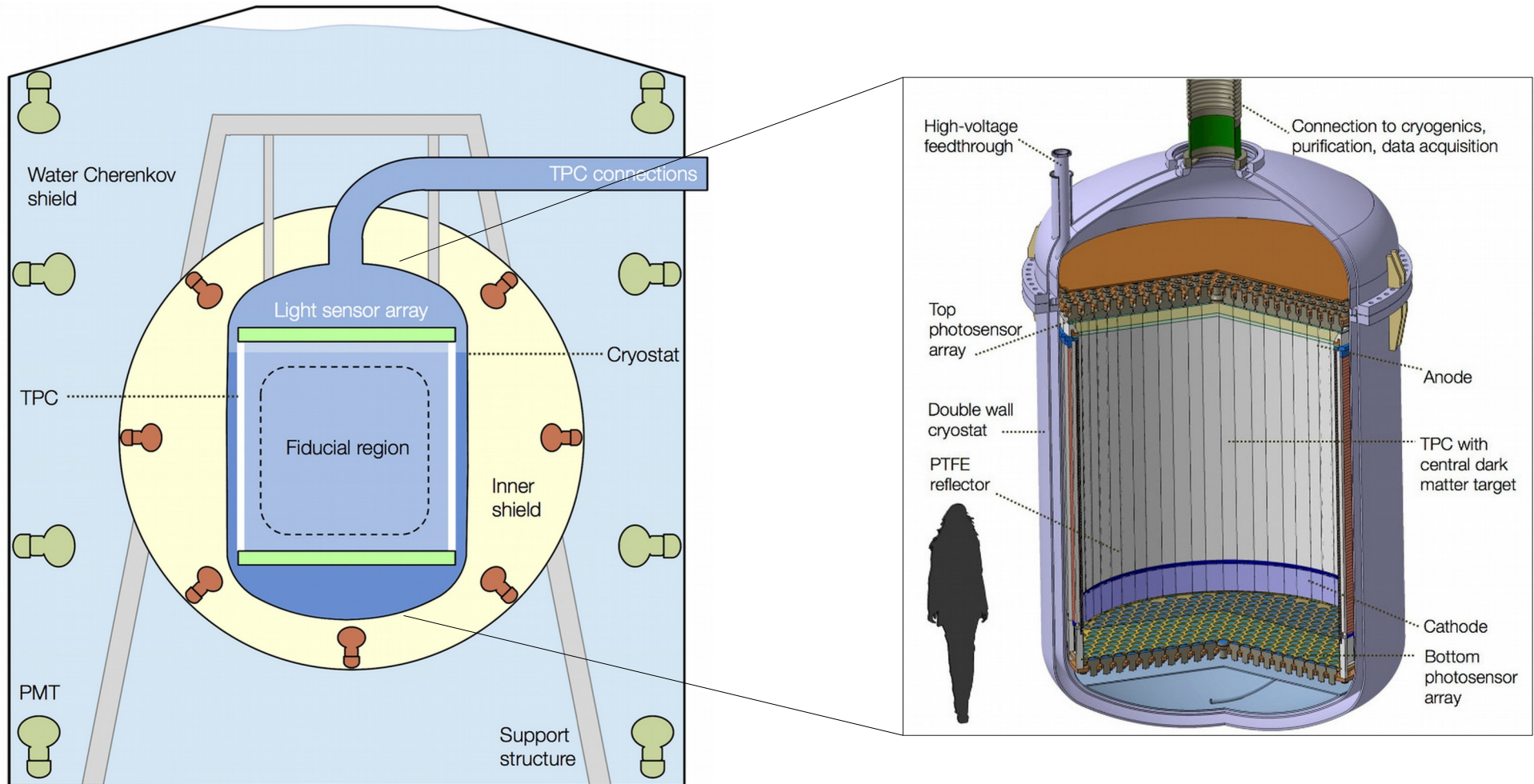
1993 layout



# Complex Detector Systems

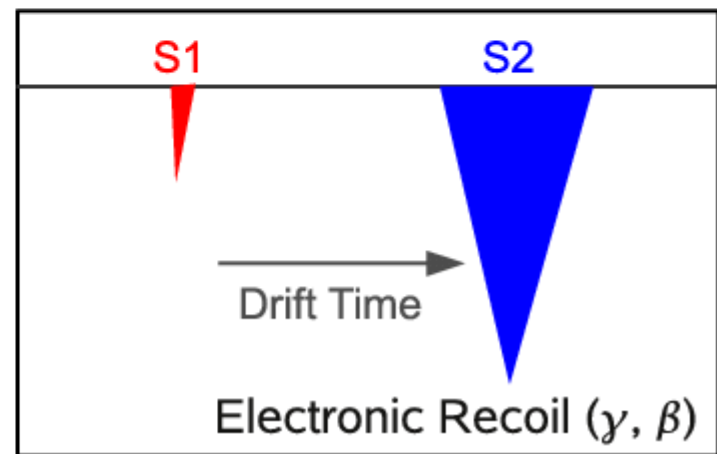
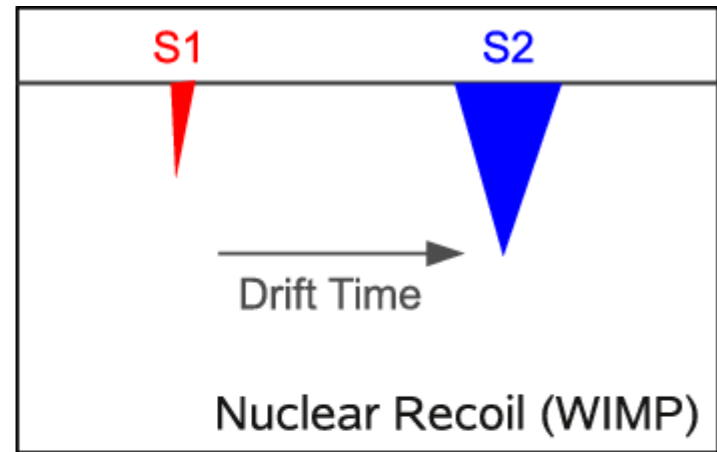
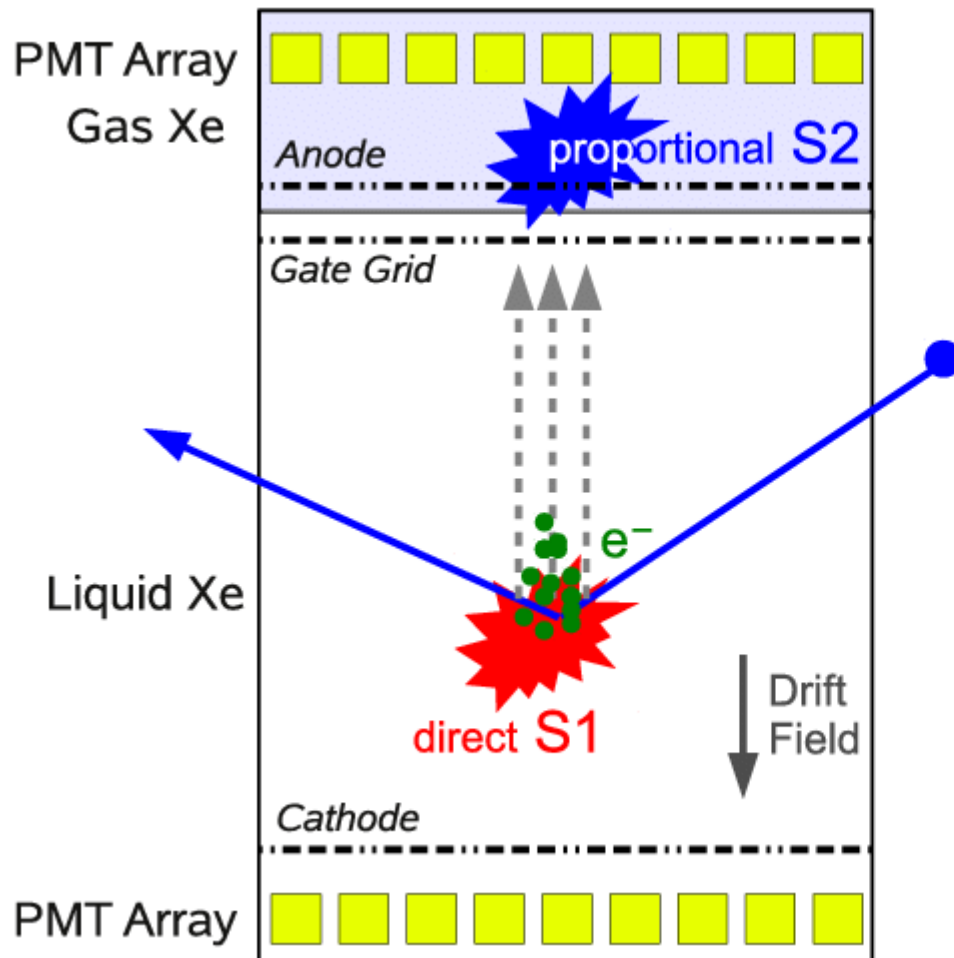


# DARWIN (DARk matter WImp search with liquid xenON)

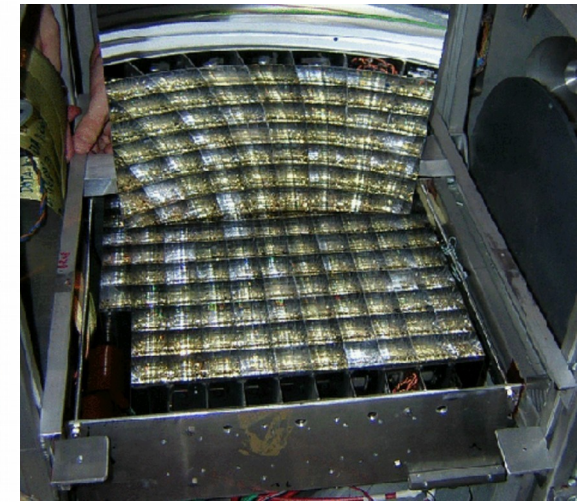
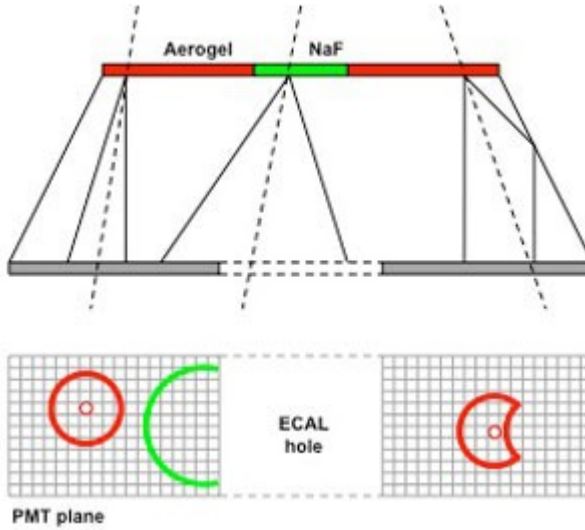




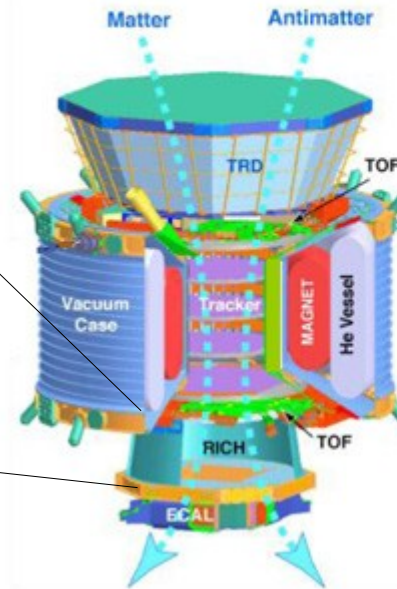
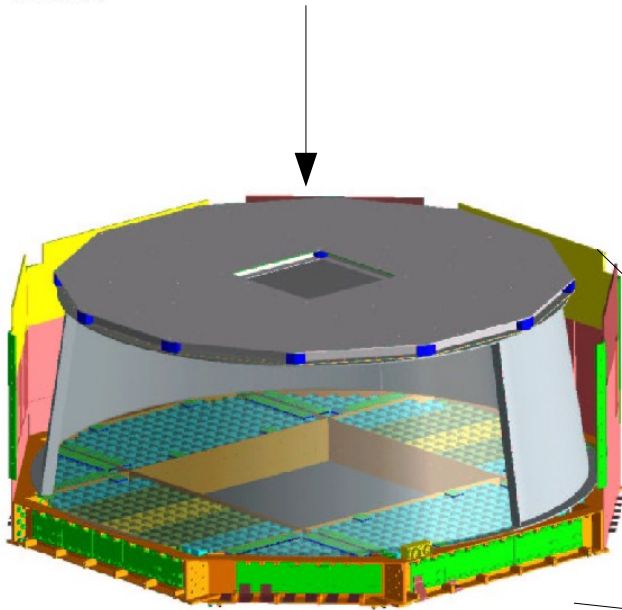
# Liquid Xenon Detectors



# AMS-2 alpha magnetic spectrometer



## AMS: A TeV Magnetic Spectrometer in Space

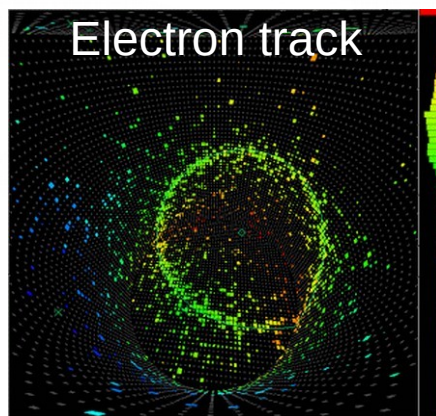
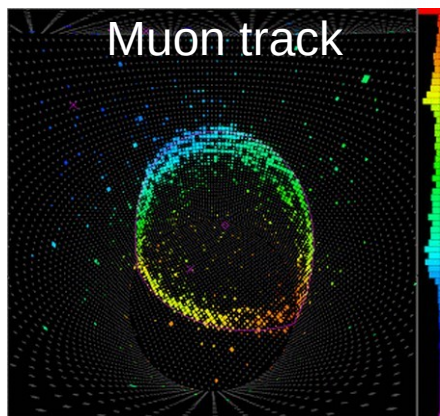
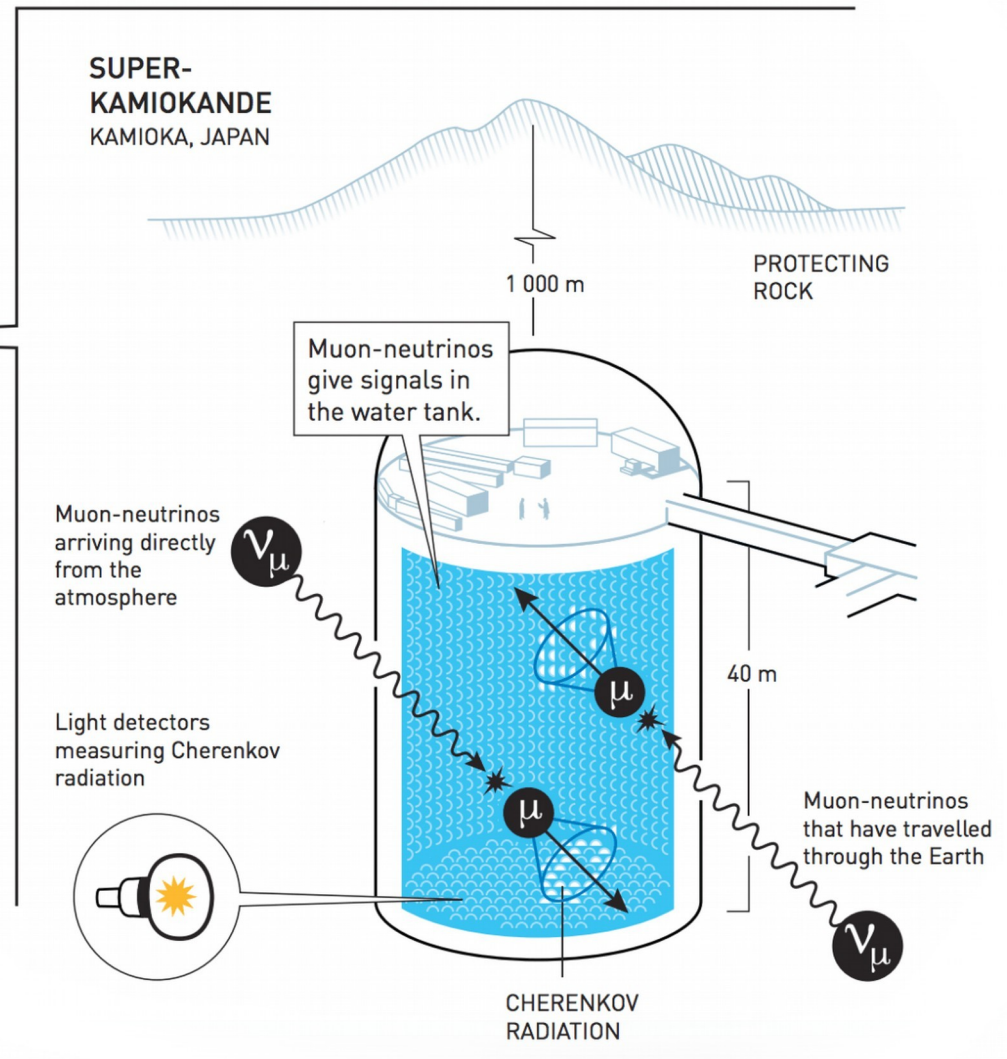
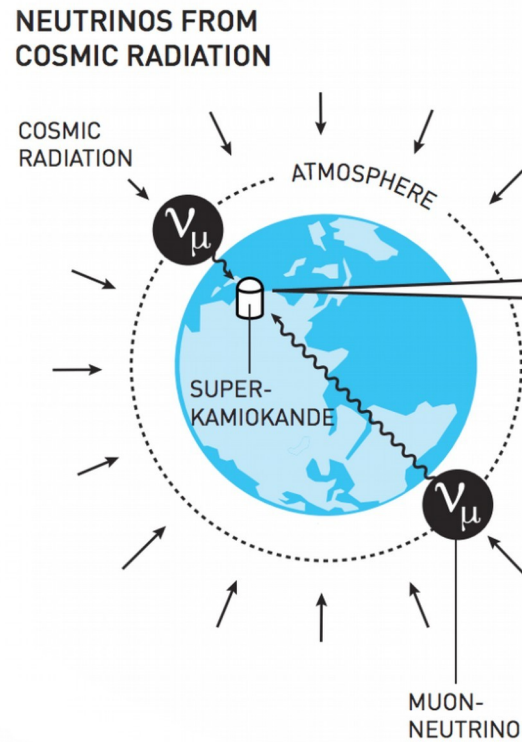


Data Signature of Various Particles in Each Detector

	$e^-$	P	Fe	$e^+$	$\bar{p}$	$\bar{He}$
TRD						
TOF						
Tracker + Magnet						
RICH						
ECAL						
Physics example	Cosmic Ray Physics Strangelets			Dark matter		Antimatter



# Super-Kamiokande neutrino detector





# CMS--detector at LHC, CERN

